

$\sin^2\theta_W(Q^2)$, Radiative Corrections and Z' Bosons

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“MOLLER at JLAB – A Special Opportunity”

Exp. Expertise & Very Clean Precise Theory

12 GeV Accelerator Upgrade

(Demands Flagship Experiments)



Outline

1. Introductory Remarks: $\sin^2\theta_W$ Status vs SM
Z Pole vs low Q^2 & BSM
2. EW Radiative Corrections to Moller A_{RL} PV
A. Czarnecki & WJM PRD (1996) Updated Uncertainty
-40±0.36% Total Reduction (*Running $\sin^2\theta_W(Q^2)$!*)
3. Z' Boson Sensitivity (extra $U(1)'$ gauge symmetry)
Strong Coupling $g' \sim 2$ $m_{Z'} \sim 8\text{TeV}$
EW GUT Coupling $g' \sim 1/2$ $m_{Z'} \sim 2\text{TeV}$
Light Z' with Very Small "Induced" Coupling
(eg Dark Parity Violation) Low Q^2 Moller Sensitivity

Introductory Remarks

$\sin^2\theta_W(Q^2)$ Status Z Pole vs low Q^2

$$\overline{MS} \sin^2\theta_W(m_Z)_{\overline{MS}} = \sin^2\theta_W^{lep} - 0.00028$$

$\sin^2\theta_W(m_Z)_{\overline{MS}}$	0.23070(26)	SLAC A_{RL}
$\sin^2\theta_W(m_Z)_{\overline{MS}}$	0.23193(29)	CERN $A_{FB}(bb)$

3 sigma difference? Tension?

$$\sin^2\theta_W(m_Z)_{\overline{MS}} \quad \underline{0.23125(15)} \quad \text{Z Pole Ave.}$$

$$\sin^2\theta_W(m_Z)_{\overline{MS}} = 2\sqrt{2}\pi\alpha/m_Z^2 G_\mu (1 - \Delta r'(m_t, m_H))$$

Through 2 loops

$$\Delta r'(m_t, m_H) = 0.0598(2) \rightarrow \sin^2\theta_W(m_Z)_{\overline{MS}} = \underline{0.23124(6)}$$

Outstanding Agreement Severely constrains BSM

S & T Constraints

Gauge Boson Self-Energy Loops

Experimental Averages

$$\sin^2\theta_W(m_Z)_{MS} = \underline{0.23125(15)} \quad \& \quad m_W = \underline{80.385(15)\text{GeV}}$$

Standard Model + S & T

$$\sin^2\theta_W(m_Z)_{MS} = \underline{0.23124(6)} [1 + 0.016S - 0.011T]$$

$$m_W = \underline{80.362(6)\text{GeV}} [1 - 0.0036S + 0.0056T]$$

$$S \approx 0.7T \quad (\text{from Z pole } \sin^2\theta_W(m_Z)_{MS} = \underline{0.23125(15)})$$

$$S = 0.07(8) \quad T = 0.10(11) \quad \Delta N_{\text{doublets}} = 2(2)$$

Little (but some) room available for “New Physics”

Constrains: New Dynamics (Technicolor),

4th Generation, SUSY, Z', Z'' (mixing)...

Best Off Z Resonance Measurements of $\sin^2\theta_W$

(Not Competitive with Z Pole)

Reaction	$\sin^2\theta_W(m_Z)_{MS}$	$\langle Q \rangle$
Cs APV	0.2283(20)	2.5MeV
E158 ee	0.2329(13)	160MeV
Q_{weak} ep	0.2329(50)	160MeV
6GeV Dis eD	0.2299(43)	1.5GeV
NuTeV $\nu_\mu N$	0.2356(16)	3-4GeV

NuTeV $\sin^2\theta_W(m_Z)_{MS} = \underline{0.2356(16)}$ (2+ sigma High)

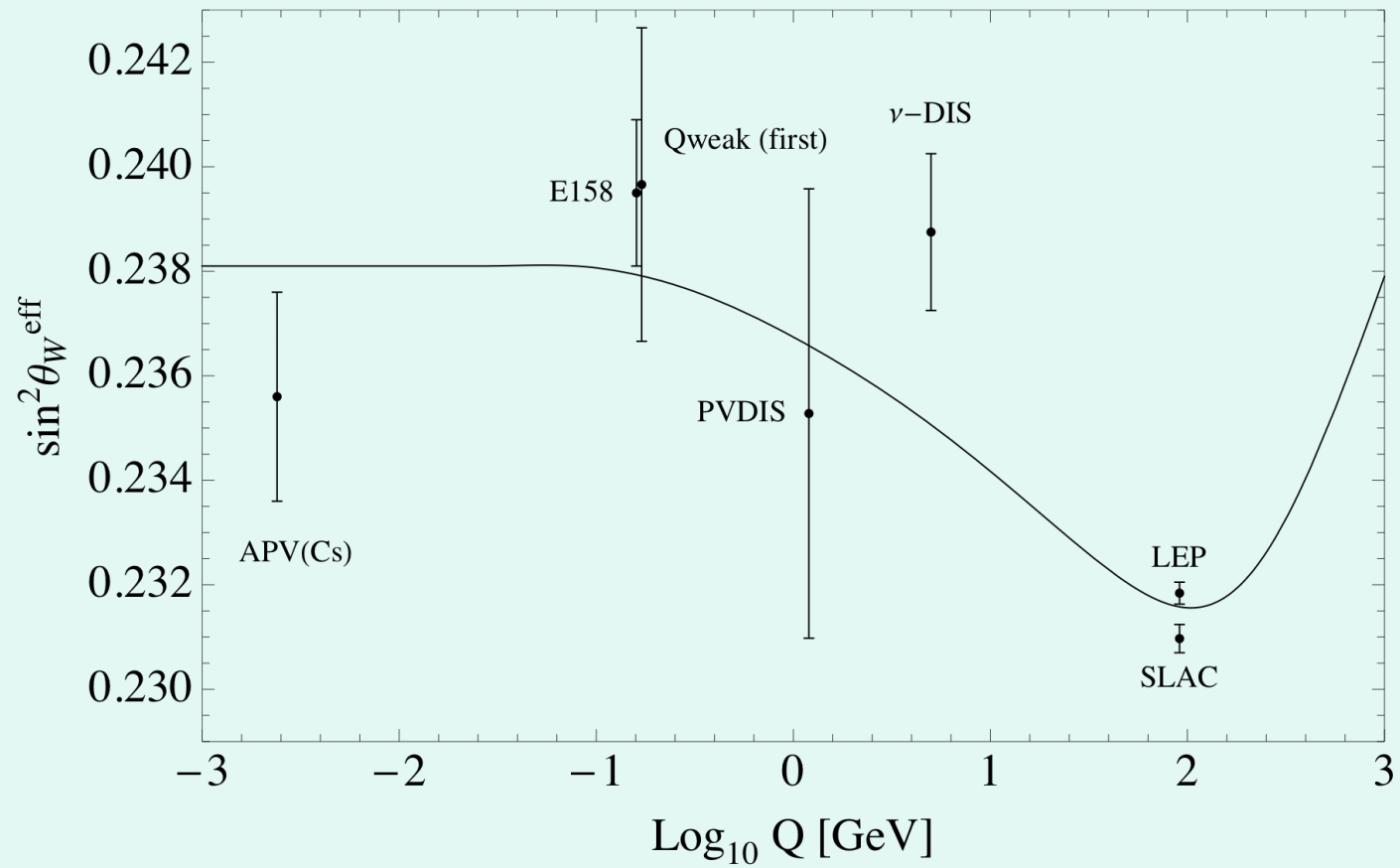
Beware The Theory Uncertainties!!!!

$A_{RL}(ee) = -131(14)(10) \times 10^{-9} \propto (1-4\sin^2\theta_W)$

Best Low Q^2 Determination $\sin^2\theta_W(m_Z)_{MS} = \underline{0.2329(13)}$

Very Clean Theory

Measurements of running $\sin^2\theta_W(Q^2)$



E158 at SLAC Pol ee→ee Moller
 $E_e \approx 50\text{GeV}$ on fixed target, $Q^2 = 0.02\text{GeV}^2$

$$A_{\text{RL}}(\text{ee}) = -131(14)(10) \times 10^{-9} \propto (1 - 4\sin^2\theta_W)$$

EW Radiative Corrections $\sim -40(3)\%$! (Czarnecki & WJM 1996)

More $\sin^2\theta_W$ Sensitivity!

Measured to $\pm 12\%$ $\rightarrow \sin^2\theta_W$ to $\pm 0.6\%$ (20 to 1)

$\rightarrow \sin^2\theta_W(m_Z)_{\text{MS}} = \underline{0.2329(13)}$ slightly high

Best Low Q^2 Determination of $\sin^2\theta_W$

$$A_{\text{RL}}(\text{ee})^{\text{exp}} = A_{\text{RL}}(\text{ee})^{\text{SM}} (1 + 0.25T - 0.34S + 7(m_Z/m_{Z\chi})^2 \dots)$$

Constrains “New Physics” eq $m_{Z\chi} > 0.6\text{TeV}$, H^\pm, S , Anapole Moment, 4 Fermion Operators ...

Together APV(Cs) & E158 $\rightarrow \sin^2\theta_W(Q^2)$ running

$\sin^2\theta_W(m_Z)_{\text{MS}} = \underline{0.232(1)}$ Good agreement with Z Pole

Goals of Future Low Q^2 Experiments

- High Precision: $\Delta \sin^2 \theta_W \sim \pm 0.0002 - 0.0004$ or better
Low Q^2 Sensitivity to “New Physics” (eg Dark PV)

SUSY Loops Sensitivity in Moller vs SM & Z pole
(Ramsey-Musolf Talk)

$S \approx 0.7T$?, Confirm Z Pole A_{ve}/m_W determination?

Z pole implies $|0.25T - 0.34S| \leq 10^{-3}$; **TRUE?**

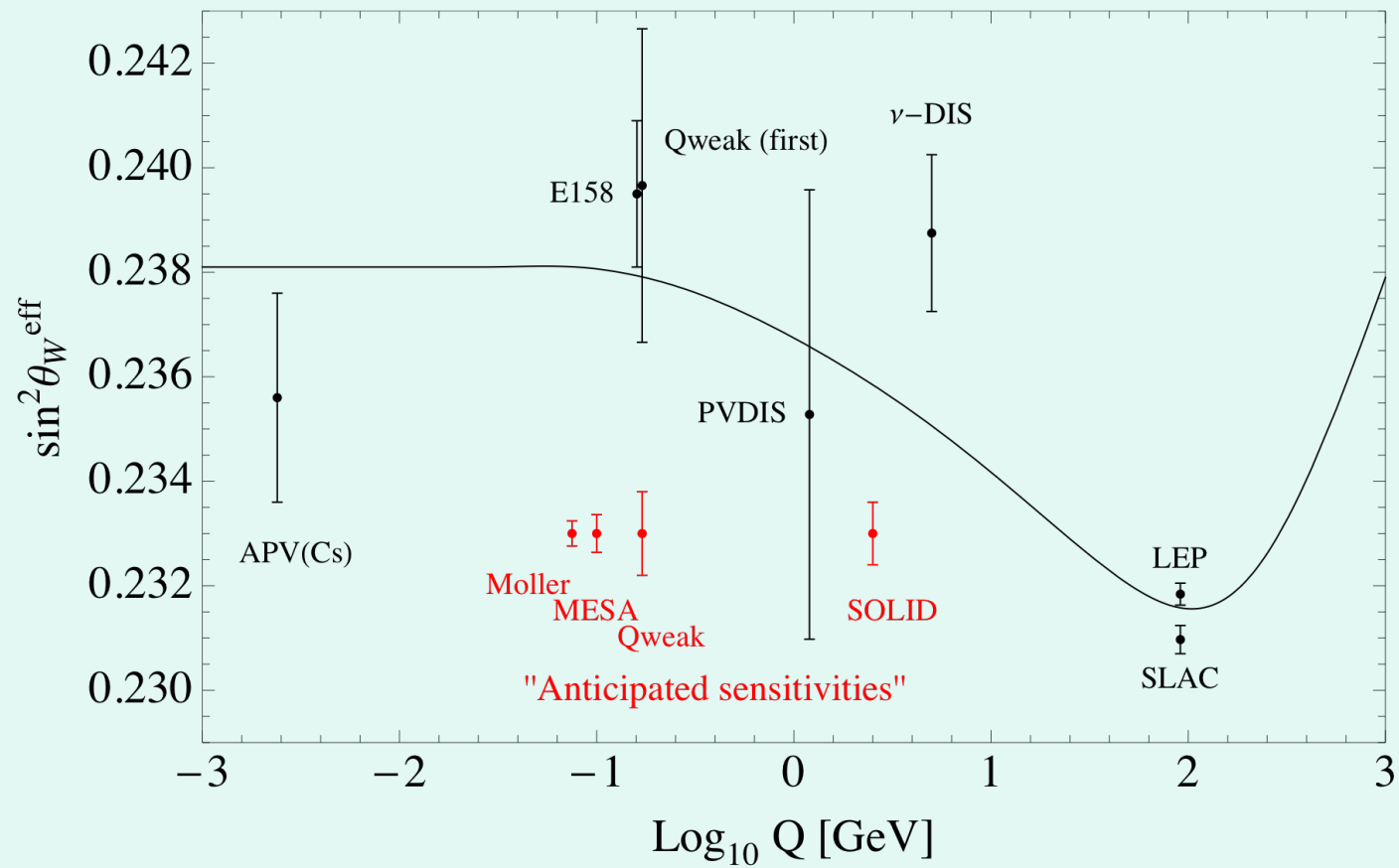
If so, Charge radius $X(Q^2)$, Z' ... of prime importance

Heavy m_Z $> 1-8 \text{ TeV}$, Model Dependent Sensitivities

(Do not interfere at Z Pole)

***Light m_Z** , $< 10 \text{ GeV}$ (Dark Parity Violation) *Sensitivity*

$\sin^2\theta_W(Q^2)$ Measurements & expected Future Sensitivities



Polarized ee, ep Asymmetries

- $A_{RL} = \sigma_R - \sigma_L / \sigma_R + \sigma_L$ Parity Violating $\propto Q^2$ very small

Experiment	$\langle Q \rangle$ MeV	$\Delta \sin^2 \theta_W$	Measurement
E158 SLAC	160	± 0.0013	ee Completed
Q_{weak} JLAB	160	± 0.0008	ep in analysis
<u>Moller JLAB</u>	75	± 0.00027	<u>ee approved</u>
MESA (ep) P2	50-100?	± 0.00037	ep Low Energy

Also improve: APV, DISeD, Neutrino Scattering...

Polarized Moller at JLAB

After 12GeV Upgrade (4 x lower E & Q²)

$A_{RL}(ee \rightarrow ee)$ to $\pm 2.4\%$

$\Delta \sin^2 \theta_W(m_Z)_{MS} = \pm 0.00027!$

Comparable to Z pole studies!

$A_{RL}(ee)^{exp} = A_{RL}(ee)^{SM} (1 + 0.25T - 0.34S + 7(m_Z/m_{Z\chi})^2 \dots)$

Explores $m_{Z\chi} \rightarrow 2\text{TeV}$ Better than APV, S=0.7T etc.

Future JLAB Flagship Experiment!

Unique Opportunity

Standard for Future Proposals

EW Radiative Corrections to Moller A_{RL} (-40%)

A. Czarnecki & WJM PRD(1996)

- $A_{RL}(ee) = \alpha (1 - 4\sin^2\theta_W)$ $\sin^2\theta_W(m_Z)_{MS} = \underline{0.23124(6)}$ or
running + $3.01(25)_{\text{hadronic}}\%$
 $\sin^2\theta_W(Q=0) = \underline{0.23820(60)}$
+ WWbox (+3.6%) γZ box...(-5.5%) partial cancellation
+ other small 1 loop corrections \rightarrow -40(3)% reduction!
E158 $\Delta A_{RL}/A_{RL} = \pm 12.5\%$ vs Running unc. $\pm 6\%$?

Erler & Ramsey-Musolf \rightarrow factor of 8.6 error reduction!

+3.01(25)% \rightarrow +2.99(3)% Theory $\pm 0.6\%$ vs Moller exp $\pm 2.4\%$

$\Delta \sin^2\theta_W^{RC} \sim \pm 0.00007!$ Pristine

Potentially another factor of 2 reduction via lattice

1 loop contributions to $\sin^2\theta_W(Q^2)$ running

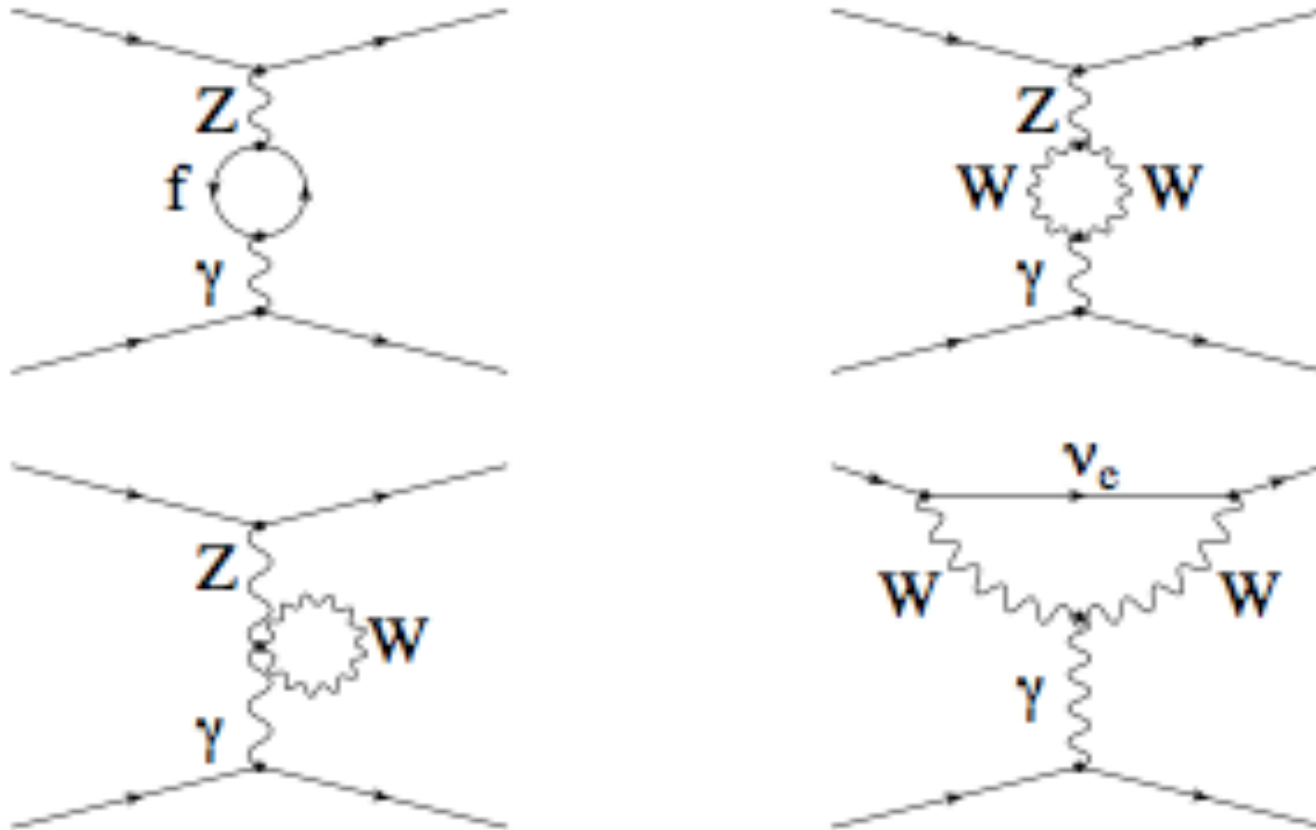


Fig. 2. $\gamma - Z$ mixing diagrams and W -loop contribution to the anapole moment.

Box Diagrams (tend to cancel) →
 very small 2 loop uncertainty ($\Delta A_{RL}/A_{RL} \sim \pm 0.3\%$)

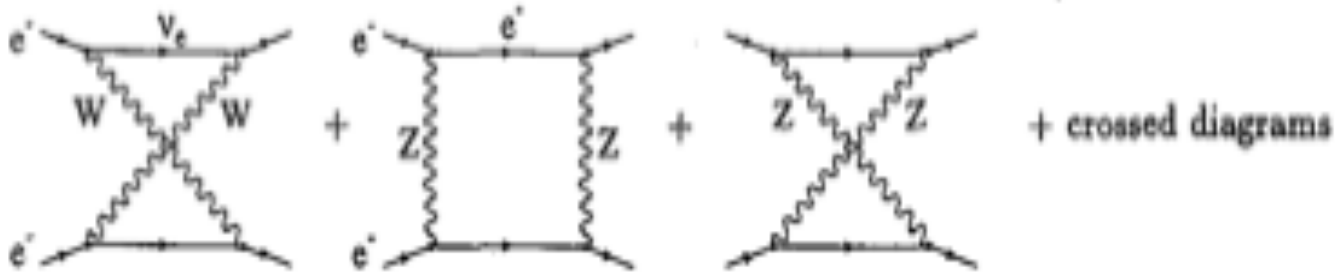


FIG. 3. Box diagrams with two heavy bosons.

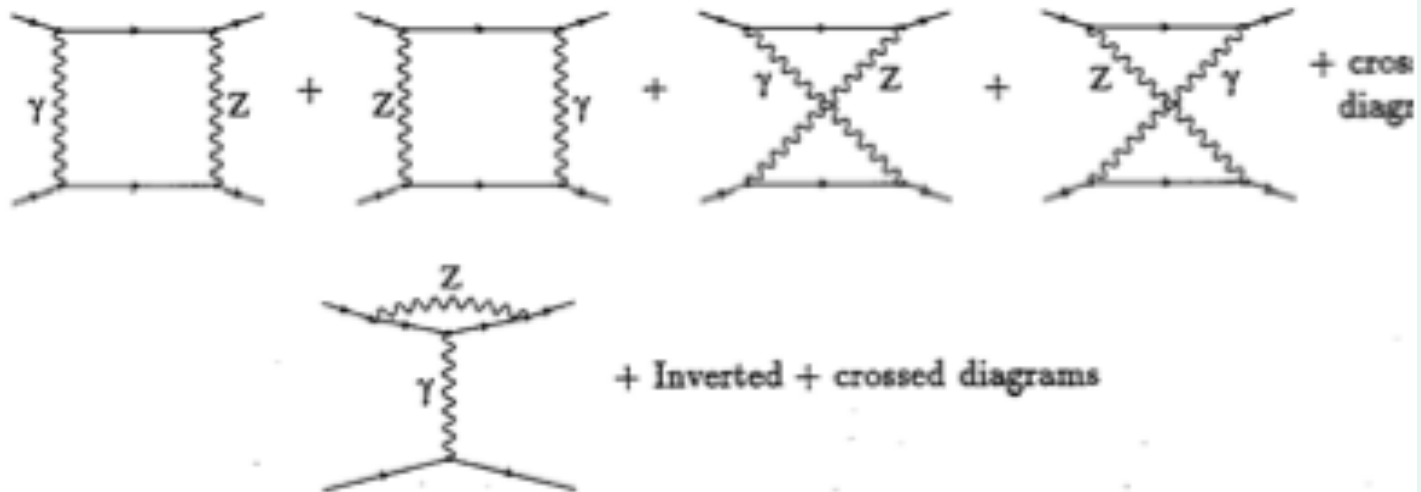


FIG. 4. Boxes containing one photon and Z-loop contribution to the anapole moment.

Z' Boson Sensitivity (extra U(1)' gauge symmetry)

Consequences of $A_{RL}(ee \rightarrow ee)$ to $\pm 2.4\%$

$$\Delta \sin^2 \theta_W(m_Z)_{MS} = \pm 0.00027!$$

$$A_{RL}(ee)^{\text{exp}} = A_{RL}(ee)^{\text{SM}} (1 + 0.25T - 0.34S + 7(m_Z/m_{Z\chi})^2 \dots \\ + 0.7R(0) + \text{SUSY} + H^- \dots)$$

Z pole $\sin^2 \theta_W(m_Z)_{MS} \rightarrow 0.25T - 0.34S$ very small. *Is it?*

SM agreement significantly constrains BSM

Deviation Implies New Physics (Many Z' Examples)

Complements LHC Z' Direct Discovery 1-8TeV

(LHC Requires relatively large Z' production & I^+I^- BR)

(Potentially) Watered down by BSM decay modes

Examples

$U(1)_\chi, U(1)_\psi$ of E_6 GUT (Mix) $\leq 2\text{TeV}$ sensitivity

Much Stronger coupling \rightarrow Better Sensitivity $\rightarrow O(8\text{TeV})$

$U(1)_{B-L}, U(1)_{L_e-L_\mu}, U(1)_{SUSY}\dots$ Many Heavy Z' Examples

LHC Sensitivity may be diluted by reduced I^+I^- BRs

Light Z' bosons with very weak SM couplings $\sim 10^{-3}-10^{-6}g$

I will illustrate **dark Z** model of **(DAVOUDIASL, LEE, MARCIANO)**

GENERALIZATION OF DARK PHOTON MODEL

The Dark Boson – A Portal to Dark Matter

- What if some $U(1)_d$ gauge symmetry from the Dark or some Other Sector contains a “Light” *Dark Photon, U Boson, Hidden Boson... Dark Z (Z_d)*

- Introduced for:
- 1) Sommerfeld Enhancement $D+D \rightarrow Z_d+Z_d$
 - 2) $Z_d \rightarrow e^+e^-$ (source of positrons, *γ -rays*)
 - 3) Cosmic Dark Matter Stability via global $U(1)_d$
 - *4) Light Dark Matter Abundance
 - *5) *Muon Anomalous Magnetic Moment*

Can we find direct evidence for such a light boson in the laboratory?

Dark Photon & Dark Z

Interacts with our particle world via

1) Kinetic Mixing $U(1)_Y \times U(1)_d$ $\epsilon e Z_d^\mu J_\mu^{\text{em}}$ $\epsilon \approx \alpha/\pi \approx 2 \times 10^{-3}$
 $\Delta a_\mu = a_\mu^{\text{exp}} - a_\mu^{\text{SM}} = 288(80) \times 10^{-11}$ (***3.6 σ discrepancy!***)
 $\approx \frac{1}{4}(\alpha/\pi)\epsilon^2$ $m_{Z_d} \approx 20\text{-}300\text{MeV}$ (see figure)

2) Z-Z_d Mass Mixing: $\epsilon_Z g/2 \cos\theta_W Z_d^\mu J_\mu^{\text{NC}}$ $\epsilon_Z = m_{Z_d}/m_Z \delta$ $\delta \approx 10^{-3}$ ($\sim v_1^2/v_2 v_s$)
Induced by extended Higgs (2 doublets + sing.) Portal
Rare Higgs $\rightarrow ZZ_d$, $K \rightarrow \pi Z_d$ & $B \rightarrow K Z_d$ decays $\sim \delta^2$,
Dark Parity Violation (probes $\delta \approx 2 \times 10^{-3}$ $\epsilon \approx 2 \times 10^{-3}$)
(DAVOUDIASL, LEE, MARCIANO) *Enhanced Phenomenology*

3) Small direct coupling eg. $U(1)_{L_e-L_\mu}$ $\sim \text{few } \times 10^{-3} e$

Example

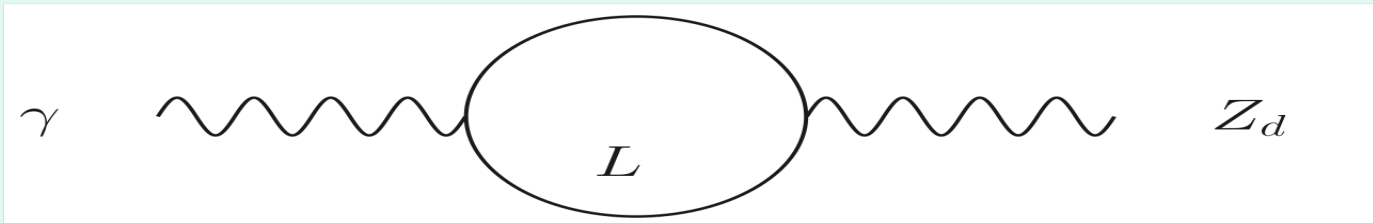
One Loop gamma- γ_d Kinetic Mixing

$$\varepsilon F_{\mu\nu} D^{\mu\nu}$$

(eg Through Heavy Charged Leptons)

That also carry $U(1)_d$ charge

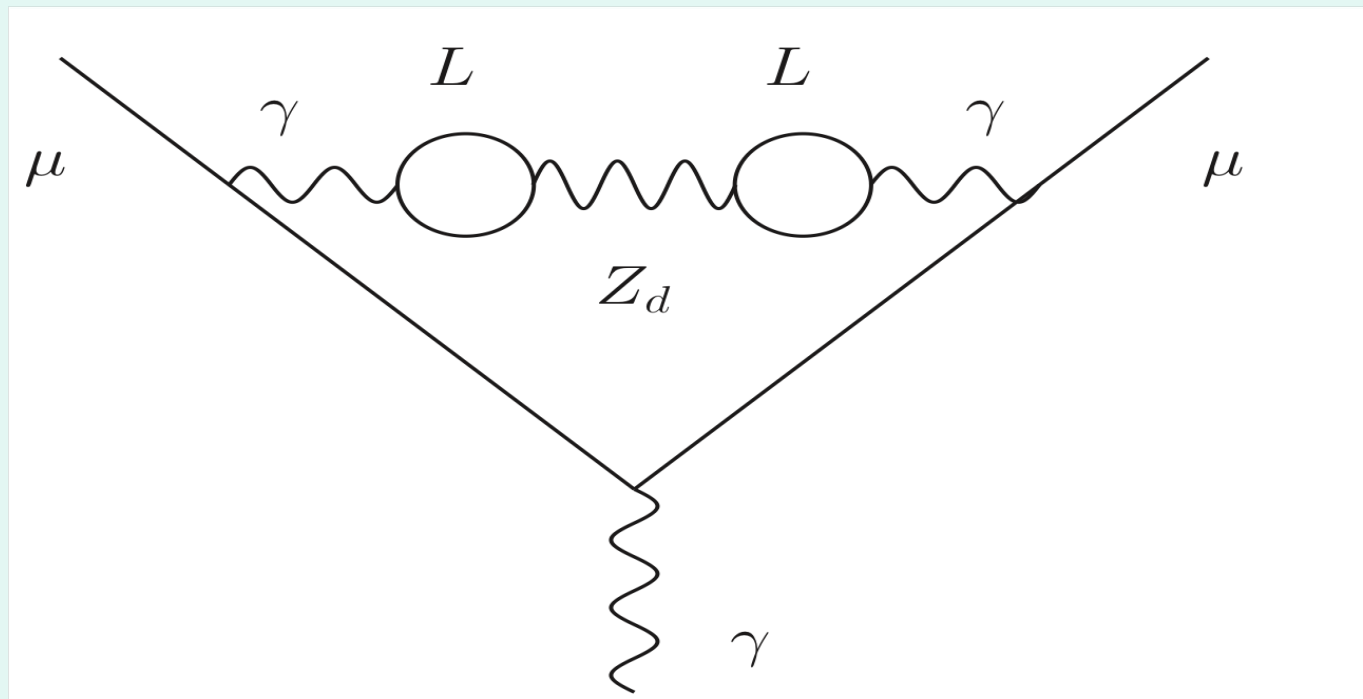
Expect $\varepsilon \sim e g_d Q Q_d / 8\pi^2 \leq O(10^{-3})$



Muon Anomalous Magnetic Moment

$$a_{\mu}^{Z_d} = \alpha/2\pi\epsilon^2 F(m_{Z_d}/m_{\mu}), F(0)=1$$

solves $(g_{\mu}-2)/2$ discrepancy $\approx 288(80)\times 10^{-11}$
for $\epsilon^2 \approx 10^{-6}-10^{-4}$ & $m_{Z_d} \approx 10-300\text{MeV}$ (see figure)

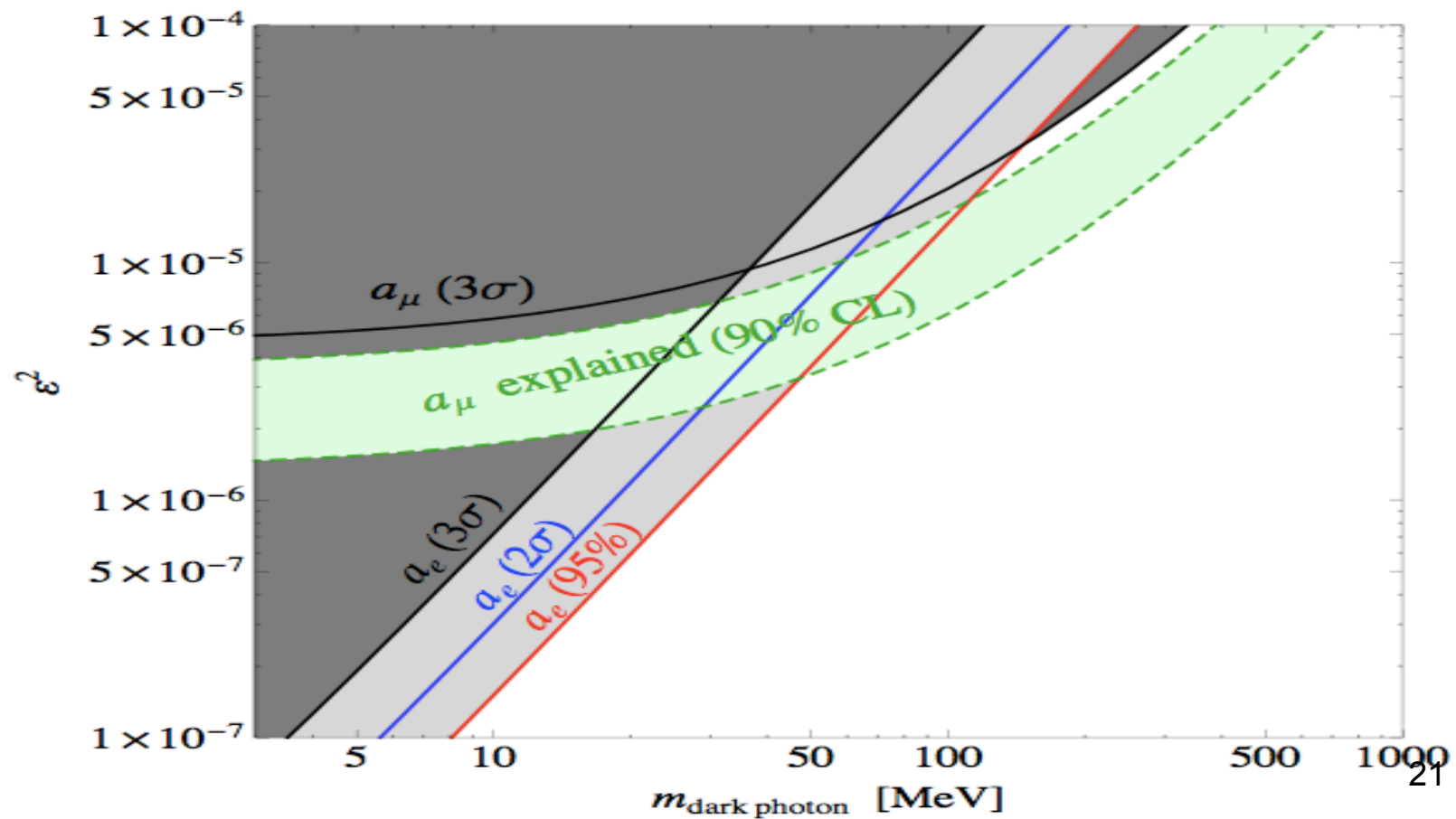


Lepton Magnetic Moment Constraints on the Dark Photon

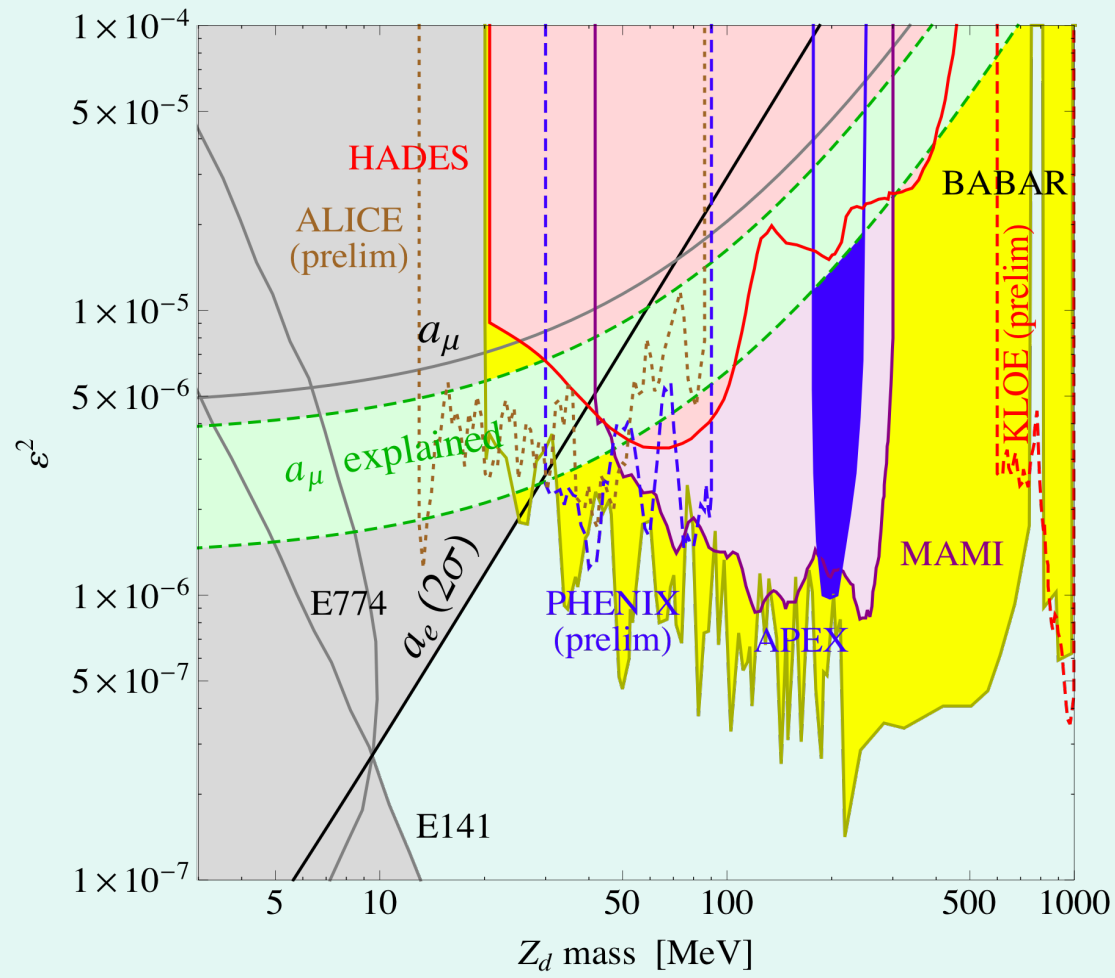
Green Band Corresponds to $a_\mu^{\text{exp}} - a_\mu^{\text{SM}} = 288(63)(49) \times 10^{-11}$ 90% CL

$g_e - 2$ Constraint DAVOUDI ASL, LEE, MARCIANO

$a_e(\text{exp}) - a_e(\text{theory}) = -1.05(0.81) \times 10^{-12}$ wrong sign!



Current Dark Photon Constraints Assuming $\text{BR}(Z_d \rightarrow e^+e^-) \sim 1$



Dark Photon/Z Muon g-2 Solution (Currently)

Recent direct searches for $Z_d \rightarrow e+e^-$

Significantly constrain Muon g-2 solution

To $m_{Z_d} \sim 20\text{-}36\text{MeV}$ (eg Recent Phenix paper), BaBar

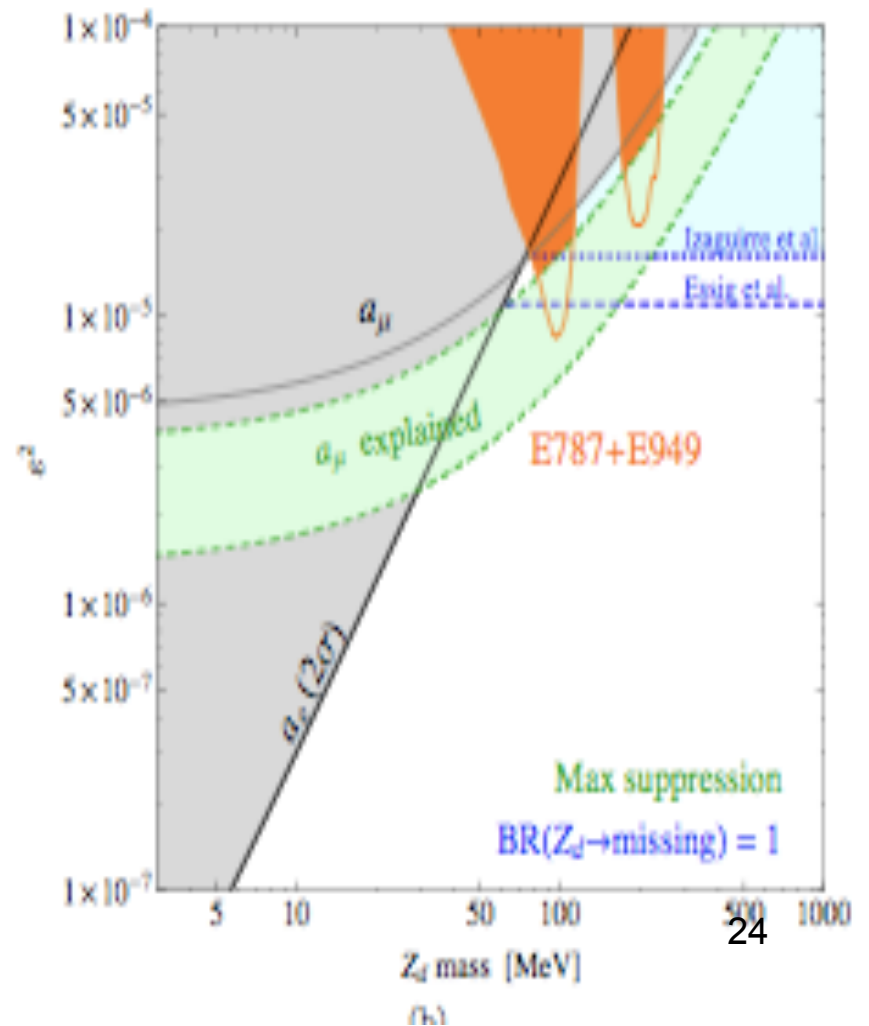
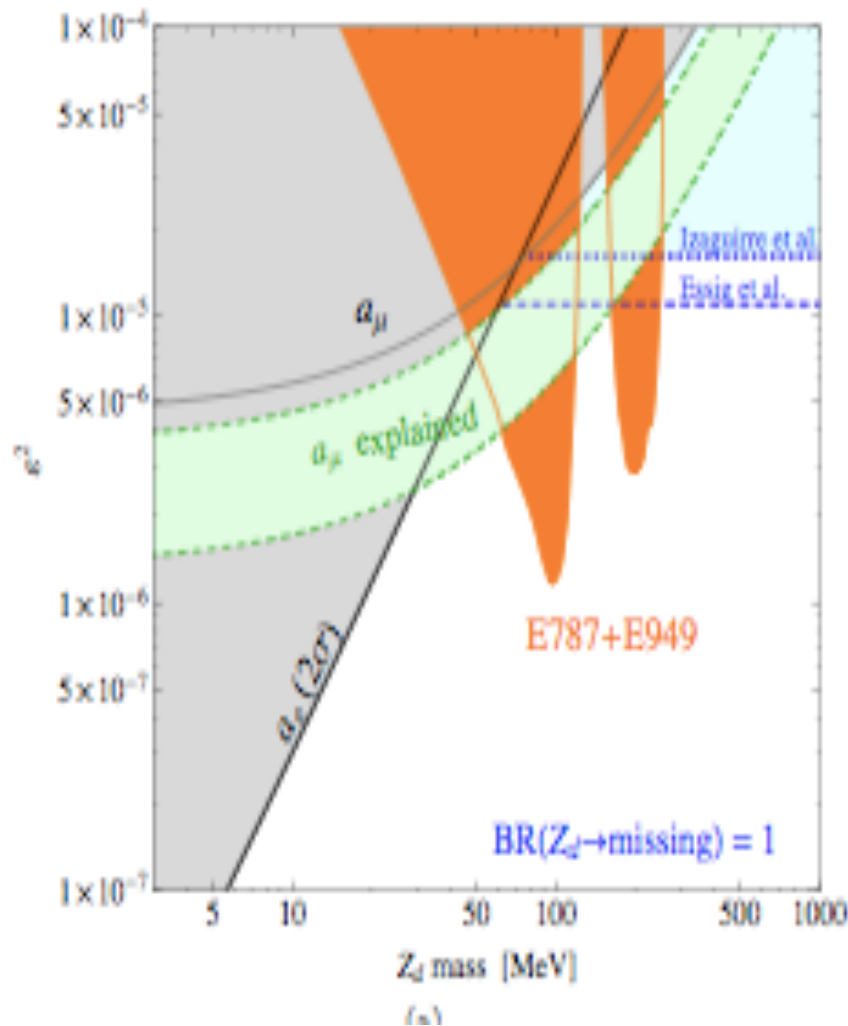
Response: Now we know where it is!

Or make $\text{BR}(Z_d \rightarrow e+e^-) \ll 1$

Introduce very light dark particle decays $Z_d \rightarrow$ missing energy

$K \rightarrow \pi^+ Z_d$ Constraints for $BR(Z_d \rightarrow \text{dark matter}) \sim 1$
 $m_{Z_d} = 100, 200 \text{ MeV}$ ruled out? $\varepsilon - \varepsilon_Z$ Interference?

H. DAVOUDIASL, H-S LEE, W. MARCIANO



Dark Parity Violation

H. DAVOUDI ASL, H-S LEE, W. MARCIANO

Effect of ε & ε_z together: (at low $Q^2 \ll m_Z^2$)

$$\Delta \sin^2 \theta_W(Q^2) = -0.42 \varepsilon \delta m_Z m_{Z_d} / (Q^2 + m_{Z_d}^2)$$

For $\delta \approx m_{Z_d} / m_Z$, $\Delta \sin^2 \theta_W(Q^2) = \pm 0.42 \varepsilon m_{Z_d}^2 / (Q^2 + m_{Z_d}^2)$

Shift largest at small $Q^2 < m_{Z_d}^2$ ($\approx O(1\%)$! Eg APV)

(1.5 sigma APV deviation) fit $\rightarrow \varepsilon \delta = 4 \times 10^{-6}$

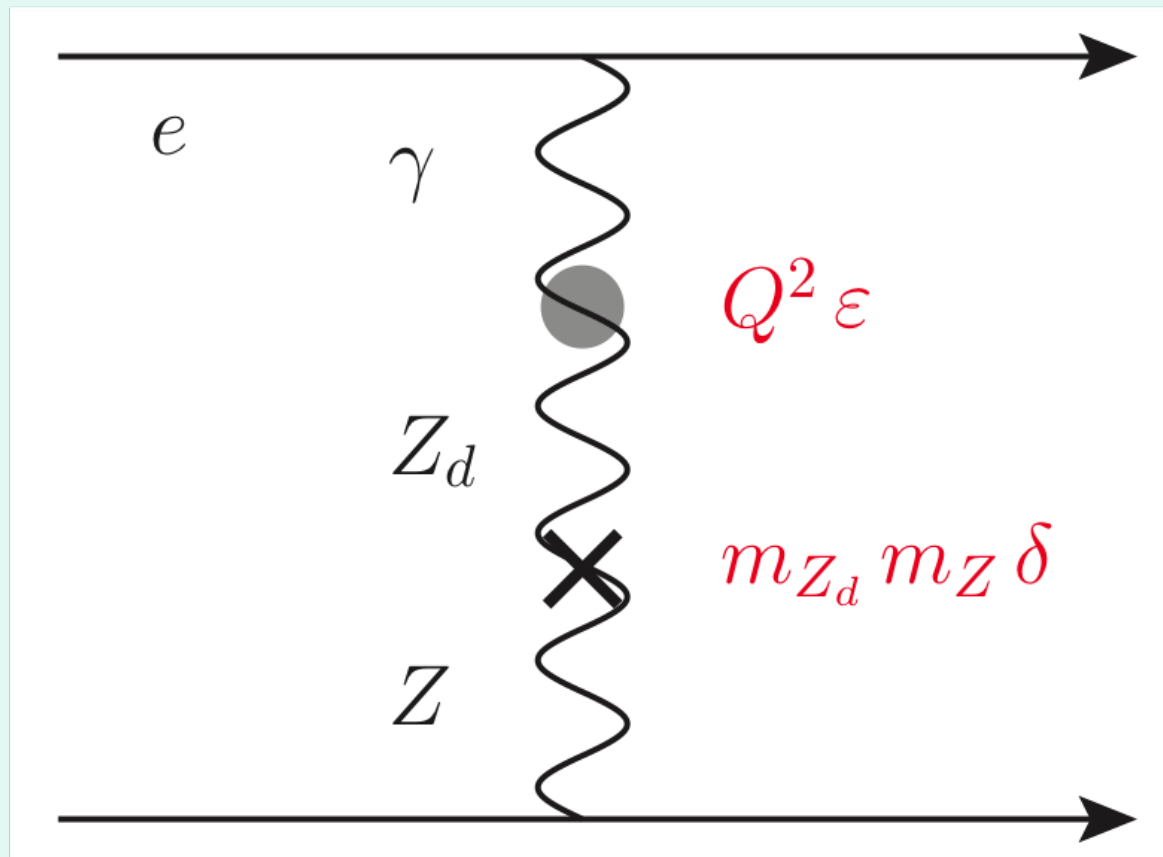
or $\varepsilon \approx \delta \approx 2 \times 10^{-3}$ for $(g_\mu - 2)$ & APV $\rightarrow m_{Z_d} \approx 50 \text{ MeV}$ region

$\sin^2 \theta_W(Q \approx 75 \text{ MeV})$ shift by $\pm O(0.5-1\%)$!!

$\varepsilon \delta$ down to $\approx 10^{-6}$ Potentially Observable

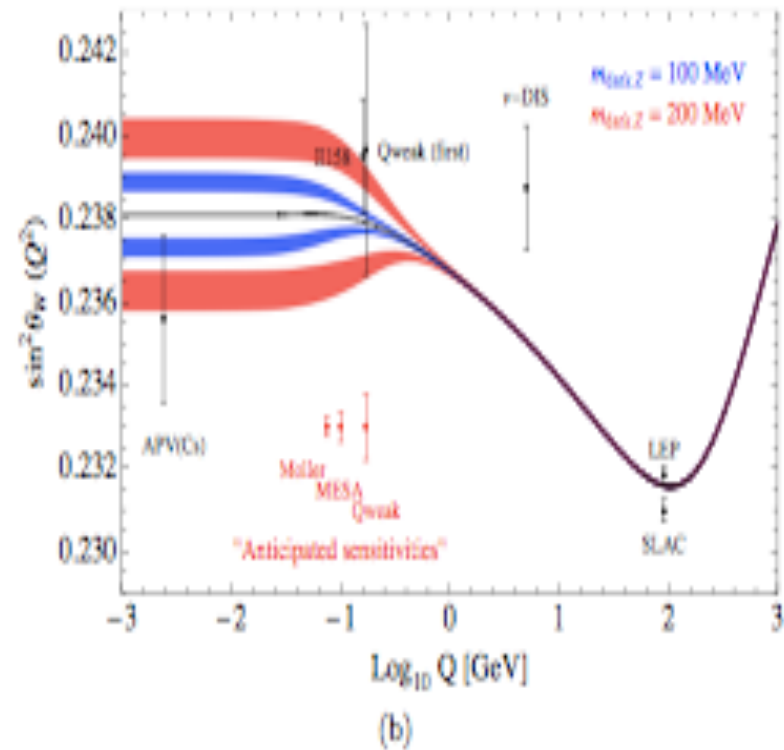
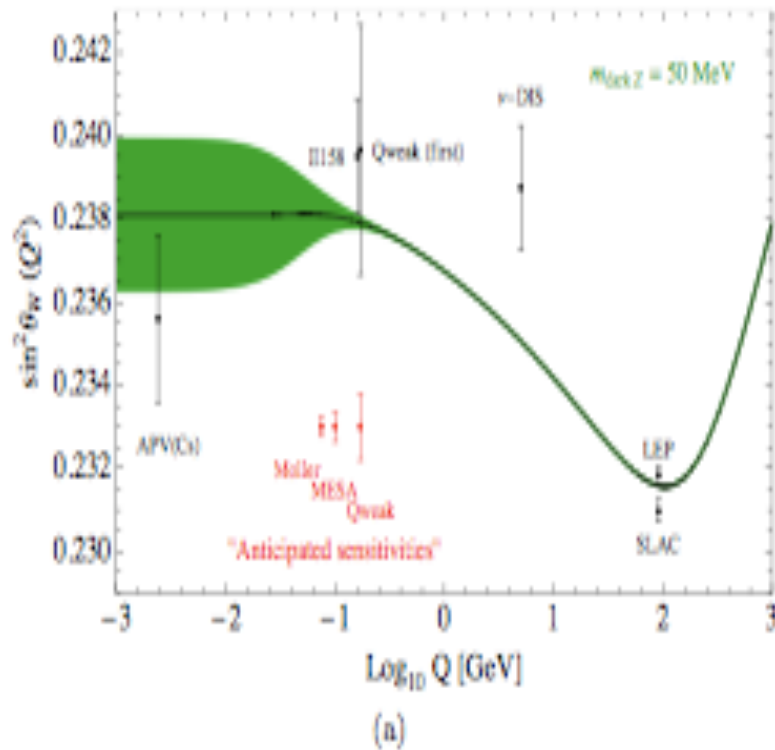
$A_{RL}(ee)$ at low Q^2 very Important

Dark Z Effect on electron scattering
Photon-Z Mixing through Z_d
from H-S Lee

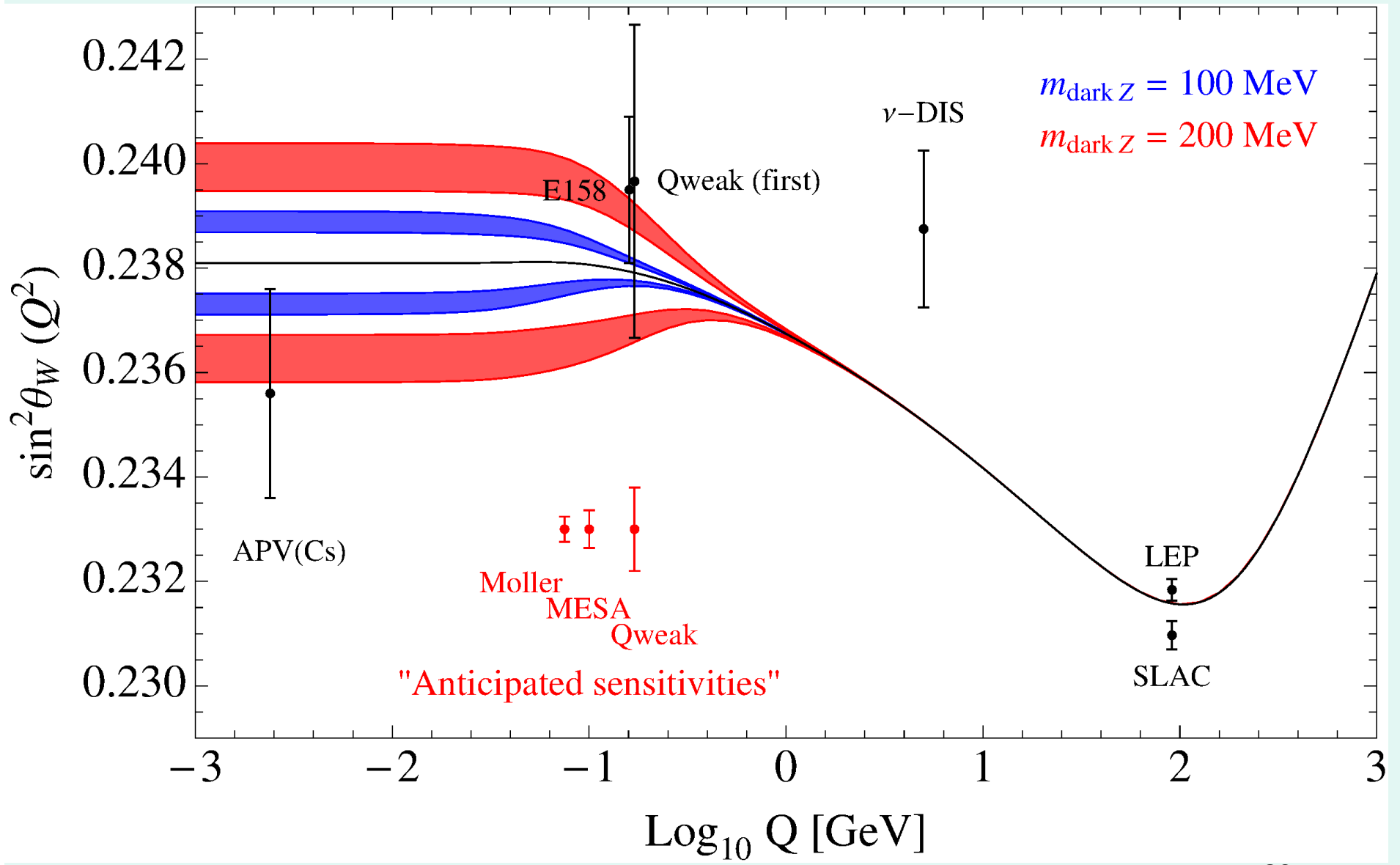


Effect of “Light” Z_d on Running

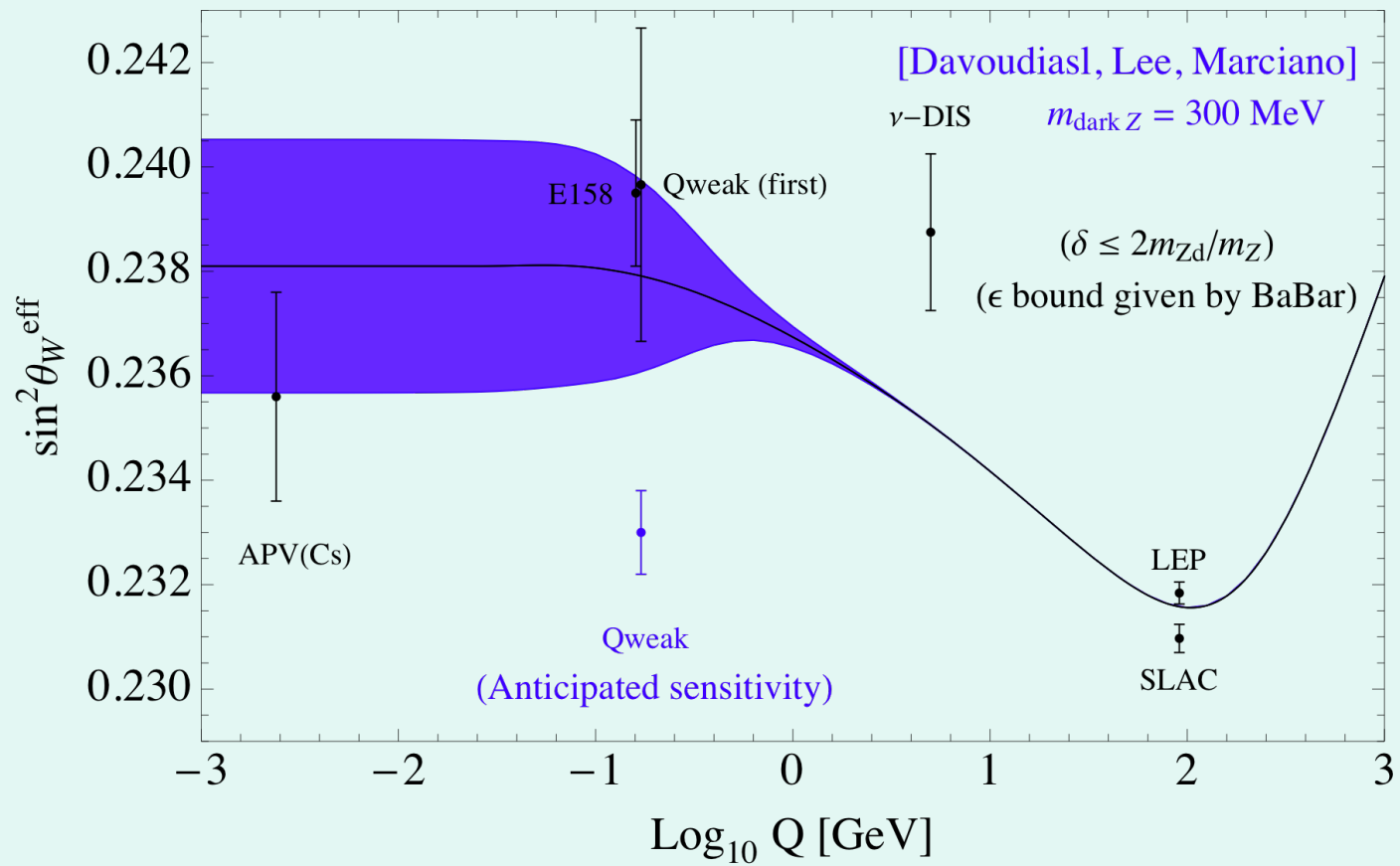
H. DAVOUDIASL, H-S LEE, W. MARCIANO



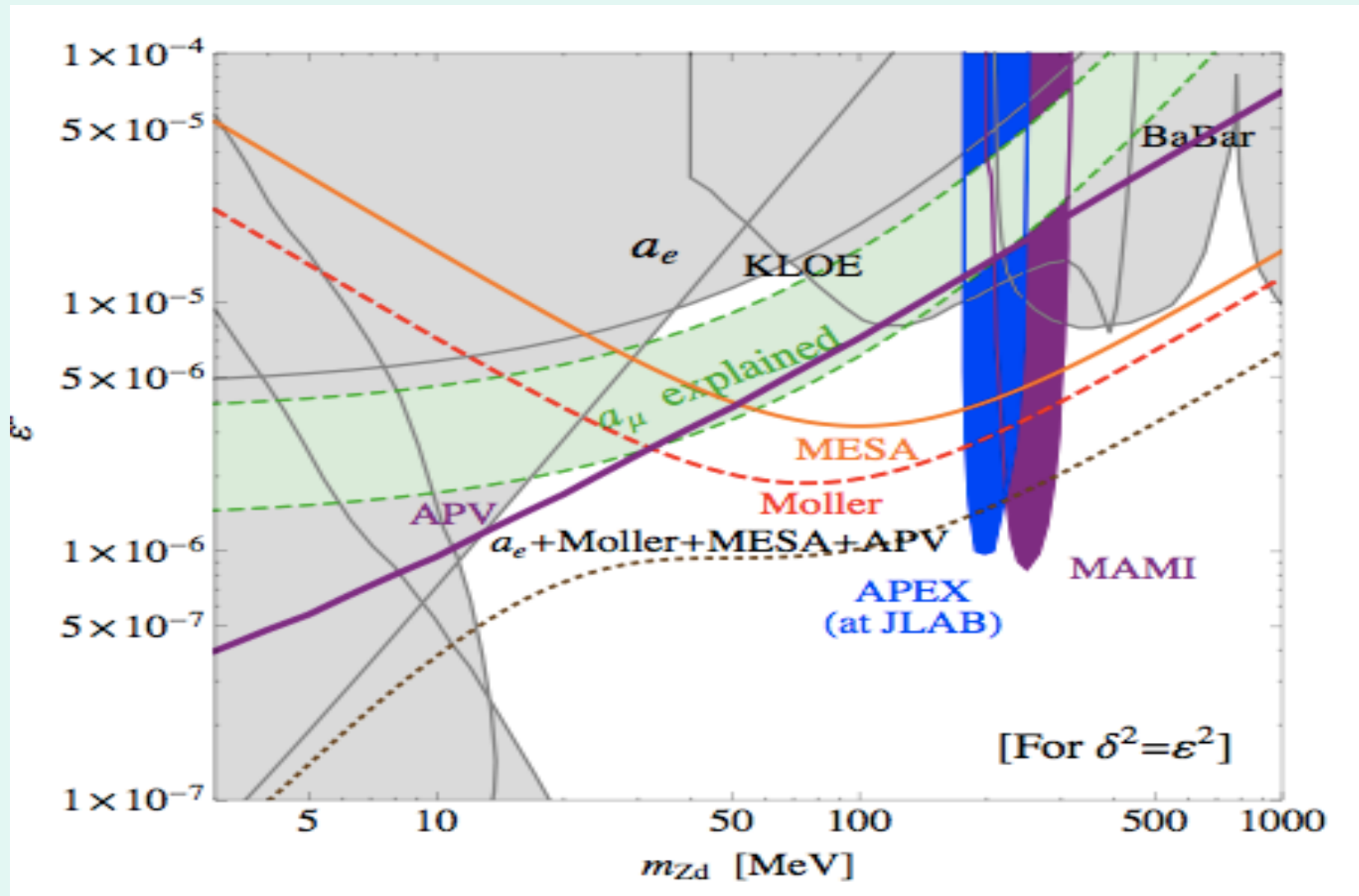
Blowup



Potential 300MeV Dark Z Effects on Running
 $0 < |\sin^2 \theta_W(0)| < 0.002$ Start to show up in Qweak



a_e & Polarized e^- Parity Violation Sensitivity (Old slide pre experimental updates)



Motivation for Moller

Very Clean, Precise, $\sin^2\theta_W(Q^2)$ Experiment at $Q=75\text{MeV}$

Factor 4 lower Q^2 relative to E158 actually improves theory!

Goal $\Delta\sin^2\theta_W(m_Z)_{MS} = \pm 0.00027!$ (Z pole competitive)

Unique hadronic $\Delta\sin^2\theta_W^{RC} \sim \pm 0.00007!$ Pristine SM Theory

No Interpretation Ambiguity!

Competitive & Complementary to Z pole & LHC measurements

Best low Q^2 opportunity for at least 10 years

“New Physics” egs. Heavy ($>1\text{TeV}$) or Light ($<10\text{GeV}$) Z’

Potential 5 sigma Discovery

JLAB Flagship - Must Do Experiment

Challenges Advance Physics (eg APV)