#### Measurement of Lepton-Lepton Electroweak Reaction MOLLER

### Introduction



# Agenda Overview & MOLLER Goals

## Krishna Kumar

Stony Brook U.

DOE Nuclear Physics MOLLER Science Review, UMass, Amherst, September 10, 2014





ENERGY



Comments on the Agenda Introduction to Møller Scattering SLAC E158 Result Introduction to MOLLER ★ Precision and Accuracy Goals ★ The Collaboration

# Thanks

Thanks to the panel members for taking the time and agreeing to participate in this review

Thanks to the DOE Office of Nuclear Physics for scheduling this review and providing the opportunity to present the physics case for MOLLER

Thanks to Michael Ramsey-Musolf and the Amherst Center for Fundamental Interactions (ACFI) for hosting this review at UMass, Amherst

# **Talk Titles and Speakers**

Parity-Violating Møller Scattering: Global Context (45): Michael Ramsey-Musolf

- Weak Mixing Angle, Radiative Corrections and New Vector Bosons (45): Bill Marciano
- Experimental Context and Overview of Technique (45): Krishna Kumar
- Overview of MOLLER Subsystems (45): Mark Pitt
- Statistics, Systematics and Run Phases (30): Kent Paschke

Impact on JLab Science Program (20): Cynthia Keppel

# Flow of Talks

### Morning Talks:

- ★ M. Ramsey-Musolf: Context for ultra-precise low Q<sup>2</sup> measurements
- ★ W. Marciano: robust theoretical prediction and unique discovery reach of low Q<sup>2</sup> weak neutral current measurements

### Afternoon Talks:

- ★ K. Kumar: Phenomenological perspective and experimental technique overview
- ★ M. Pitt: Description of Main MOLLER Subsystems
- ★ K. Paschke: Systematic Control and Progressive Run Phases
- ★ C. Keppel: The JLab Hall A Program Science Impact

# **Charge Elements**

### Scientific Significance

- ★ Morning talks + first part of K.Kumar talk
- Impact on Advancement of Nuclear Physics
  - ★ Morning talks + K.Kumar talk

### Research Effort Needed

#### ★ W. Marciano talk (theory), K.Kumar and JLab talks (experiment)

- Feasibility and First Three Year Impact
  - \* Afternoon talks, especially K. Paschke talk
- Impact of Implementation on JLab Program
  - \* Afternoon talk by C. Keppel









# The Standard Model Prediction: Remarkably Well-Known Radiative Corrections

$$A_{PV} = \frac{\rho G_F Q^2}{\sqrt{2}\pi\alpha} \frac{1-y}{1+y^4+(1-y)^4} \left\{ 1 - 4\kappa(0) \sin^2 \theta_W(m_Z)_{\overline{\text{MS}}} + \frac{\alpha(m_Z)}{4\pi\hat{s}^2} - \frac{3\alpha(m_Z)}{32\pi\hat{s}^2\hat{c}^2} (1-4\hat{s}^2)[1+(1-4\hat{s}^2)^2] + F_1(y,Q^2) + F_2(y,Q^2) \right\}$$

 $\mathbf{Q}_{\mathbf{W}}^{\mathbf{e}} = \mathbf{1} - 4\sin^2 heta_{\mathbf{W}} \sim \mathbf{0.075} \Longrightarrow \mathbf{0.045}$ 

The small size of the coupling, further reduced by radiative corrections, will be a recurring theme: it eases the pressure on "normalization" errors

# **The Standard Model Prediction: Remarkably Well-Known Radiative Corrections** $A_{PV} = \frac{\rho G_F Q^2}{\sqrt{2\pi\alpha}} \frac{1-y}{1+y^4 + (1-y)^4} \{1 - 4\kappa(0)\sin^2\theta_W(m_Z)_{\overline{\text{MS}}}\}$ + $\frac{\alpha(m_Z)}{4\pi\hat{s}^2} - \frac{3\alpha(m_Z)}{32\pi\hat{s}^2\hat{c}^2}(1-4\hat{s}^2)[1+(1-4\hat{s}^2)^2]$ + $F_1(y,Q^2) + F_2(y,Q^2)$ Czarnecki and Marciano (1995) $\mathbf{Q}_{\mathbf{W}}^{\mathbf{e}} = \mathbf{1} - 4\sin^2 \theta_{\mathbf{W}} \sim \mathbf{0.075} \Longrightarrow \mathbf{0.045}$ The small size of the coupling, further reduced by radiative corrections, will be a recurring theme: it eases the pressure

on "normalization" errors



Krishna Kumar, September 10 2014



### E158 Physics Implications 2003

Unique discovery space probed: Complementary to Colliders Limits on "New" Physics



 $A_{PV}$  = (-131 ± 14 ± 10) x 10<sup>-9</sup> Tree-level prediction: ~ -270 ppb

Introduction to MOLLER

### E158 Physics Implications 2003

Unique discovery space probed: Complementary to Colliders Limits on "New" Physics



 $A_{PV}$  = (-131 ± 14 ± 10) x 10<sup>-9</sup> Tree-level prediction: ~ -270 ppb SM with all corrections: -154 ppb

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### E158 Physics Implications 2003

Unique discovery space probed: Complementary to Colliders Limits on "New" Physics





# **Projected Uncertainty**



# **MOLLER Collaboration**

~120 Collaborators, 30 institutions, 6 countries

Expertise from several generations of successful parity experiments

#### Spokesperson: K. Kumar, Stony Brook U.

**Executive Board Chair and Deputy Spokesperson: M. Pitt, Virginia Tech** 

#### **Other Executive Board Members**

Dave Armstrong (William and Mary) Javier Gomez (JLab) Cynthia Keppel (JLab) Frank Maas (U. Mainz) Juliette Mammei (U. Manitoba) Kent Paschke (U. Virginia) Paul Souder (Syracuse U.)

#### **MOLLER Subsystem Leaders**

Polarized Source: G. Cates (U. Virginia) Beam Instrumentation: M. Pitt (Virginia Tech) Hydrogen Target: S. Covrig (JLab) Spectrometer: J. Mammei (Manitoba) Integrating Detectors: M. Gericke (Manitoba) Tracking Detectors, D. Armstrong (William and Mary) Polarimetry: K. Paschke (U. Virginia) Electronics/DAQ: R. Michaels (JLab) and P. King (Ohio) Simulations: S. Riordan (UMass) and D. McNulty (Idaho State)

If/when MOLLER is ready to move forward as a funded project, the governance structure will be appropriately expanded.

**Collaborators Present at Review D.** Armstrong (College of William and Mary) ♦ G. Cates (U. Virginia) M. Gericke (U. Manitoba) **E.** Ihloff (MIT) **K. Kumar (Stony Brook) K. Paschke (U. Virginia)** M. Pitt (Virginia Tech) **S. Riordan (UMass)** P. Souder (Syracuse U.) Krishna Kumar, September 10 20 Introduction to MOLLER