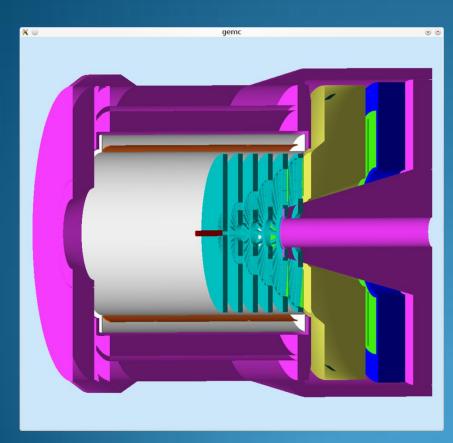
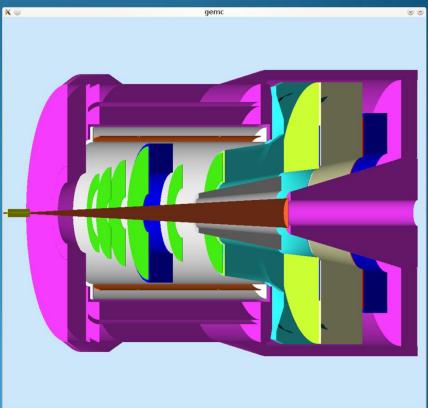
HBD -- BRIEF





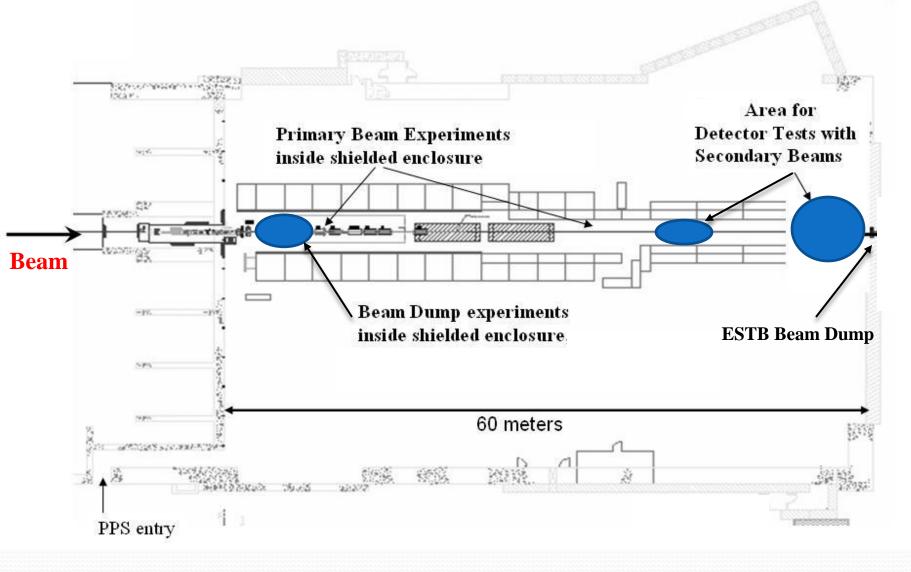
Preparation for Next Test Beam



- SLAC ESTB (End Station Test Beam)
 - Back ~1/3 of SLAC for light source.
 - 120 Hz
 - "moderate" beam
 - 5 Hz (minimum) pickoff → Target → ESA (~1 e/pulse)
- We are 1st outside users--first beam arrives Tuesday(?)
- Our run = May 1st or shortly thereafter.

Glory Days

ESA Experimental Area

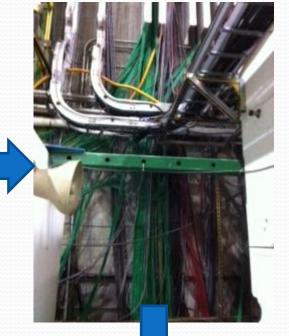


Today...Junk Pile (useful for test beam)

Flexible

Much old stuff to borrow.





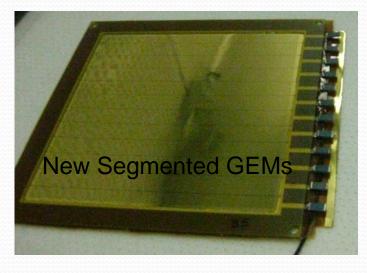
Technical staff





We're ready...





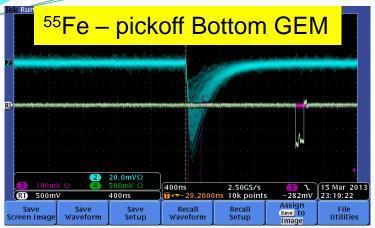


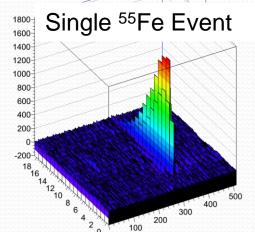




- GEM stack assembled & running w/o CsI.
- Evaporate mid > late April (effort & \$\$\$ to keep fresh)

Summary - Our Baby is Ready!



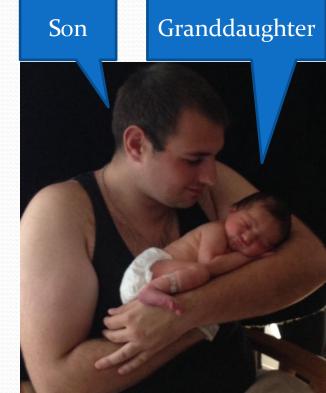


Event Display







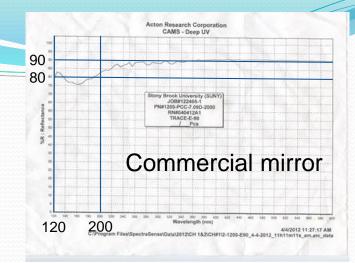




Backups

Mirror Developments

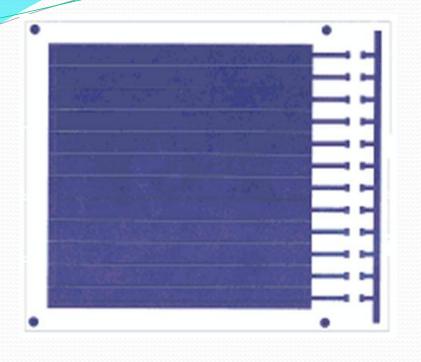
- Cherenkov photon yield primarily at small λ .
- Deep UV mirrors use 250 A^o MgF₂
 overcoat to act as dielectric mirror.
- Plans:
 - Year 1: Develop in-house manufacture of small mirrors.
 - Year 2: Scale up to use Big Mac
- First in-house mirror made!
- Reflectivity tests pending at BNL.

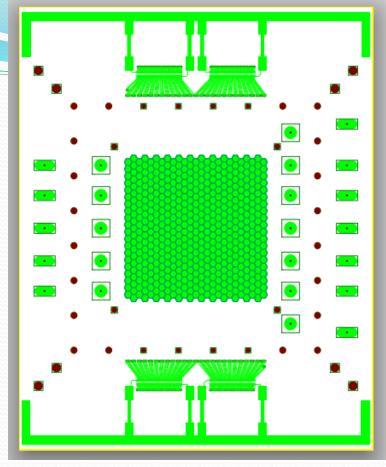






Changes for SLAC

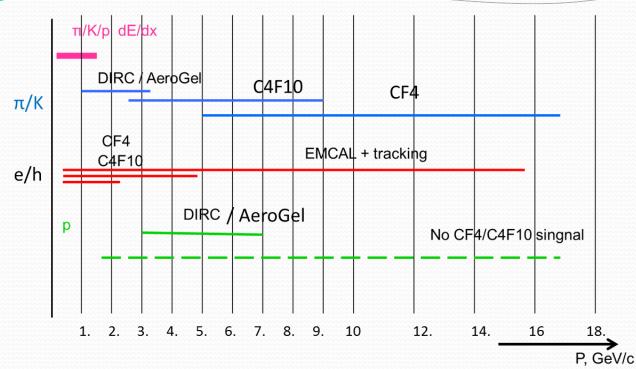




- New GEM foils...multiple strips like PHENIX.
- Second pad plane...hexagons to see rings.
- New trip detection system:
 - Capacitive coupling off resistor chain.
 - Integrated with PHENIX HBD HV relays.

Forward RICH

Electron and Hadron PID



- RICH particle ID involves a limited dynamic range of momenta set by gas index of refraction.
- The highest momenta rely on the lowest n.
- Our R&D targets the highest momenta with a CsI photo-cathode RICH.
- Major Issue: Reflectivity of mirrors deep in the UV.