Impulse Approximation limitations to the (e,e'p) reaction on <sup>208</sup>Pb Identifying correlations and relativistic effects in the nuclear medium : Analysis Report

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## **Physics Goals**

- Long Range Correlations search and spectroscopic factors
  - Measure spectroscopic factors for states near the Fermi level. Spectroscopic factors depend on short range correlations (SRC) and long range correlations (LRC).
  - Measure cross sections for these low lying states to 500 MeV/c in  $P_{miss}$ . Excess strength here is theoretically identified as due to LRC.
  - Search for  $Q^2$  dependence of spectroscopic factors
- ② Identify dynamical relativistic effects in nuclear structure. Measure cross section asymmetry A<sub>TL</sub> around the three momentum transfer. Relativistic mean field theory predicts an A<sub>TL</sub> dependence on P<sub>miss</sub> < 300(MeV/c) due to dynamical enhancement of the lower component of the nucleon wave function. Calculations which do not include the enhancement of the lower component predict a substantially different A<sub>TL</sub> behavior.

- Used doubly magic <sup>208</sup>Pb target
- Also had <sup>209</sup>Bi and <sup>12</sup>C targets.
- Measure the  ${}^{208}$ Pb $(e, e'p){}^{207}$ Tl cross sections
- Collected data at both sides of  $q: 0 \le P_{miss} \le 500 (MeV/c)$
- Had true quasi-elastic Kinematics with  $X_B \approx 1, q = 1(GeV/c), \omega = 0.433(GeV)$
- Studies of  $Q^2$  dependence on spectroscopic factors for  $P_{miss} = 0$  were performed at three  $Q^2$  points
- Had two separate runs, Run 1 on March 2007 and Run 2 on January 2008
- Run 2 had a thick and thin lead target

### Challenges to Experiment

- Lead and Bismuth melt under intense beam.
- Required high beam current for luminosity
- To prevent melting, Lead and Bismuth were sandwiched by two diamond foils.
- As a result Carbon was present in all Lead and Bismuth spectra



- Good energy resolution GEANT simulations show resolutions for a perfect spectrometer to be 1 (*MeV*) at FWHM. (to separate individual states)
- Good raster correction (<sup>208</sup>Pb and <sup>209</sup>Bi required large raster)
- Good knowledge of <sup>12</sup>C(e, e'p)<sup>11</sup>B cross sections (diamond foil in <sup>208</sup>Pb and <sup>209</sup>Bi targets)
- Good knowledge of luminosity (no hydrogen in targets to monitor luminosity)
- Good coincidence time (to get rid of random coincidences in high P<sub>miss</sub> kinematics)

### Achieved Goals: Run 1

A great expenditure of effort has been made in:

- ✓ Establish the raster correction
- ✓ Improving the coincidence time ( $\sim$  2.5ns)
- ✓ Improving the optics database

This part of the analysis is almost finished and we obtain reasonable good resolution.



# **Optics Optimization**

- Elastic  ${}^{12}C(e, e'p)$  data was used for the optics calibration
- Instead of using optimize++, a genetic algorithm, developed by Udias et al.<sup>1</sup>, was applied.
- This way all the database coefficients of a variable could be obtained at the same time.
- Genetic algorithms are more powerful than just gradient minimization alone.
- The code, written in Fortran, is, as of yet, not well documented to be shared, but hopefully will be in the future.

<sup>1</sup>C. Fernández-Ramírez, E. Moya de Guerra, A. Udías, and J. M. Udías. Phys. Rev. C **77**, 065212 (2008)

## **Optics** Optimization

- With the optimized database and with the appropriate raster correction, good resolution has been achieved.
- Though, it is finite, so we are only able to see two peaks



## Charge Normalization

- In this experiment, no Hydrogen was present to monitor the charge, so we have to be confident of the beam current incident on the target.
- Large raster hit on the target frame on some runs, and required us to compute the effective charge after cuts were applied to remove these regions.



- Due to the finite experimental acceptances, the data cannot be compared directly with theory, which is obtained from central kinematical parameters)
- As such, we used monte carlo simulations to compare the data with the theory averaged through the experimental acceptances.
- Simulations were performed with:
  - GEANT: Cross sections from factorized calculation with spectral function from Udías et al.
  - MCEEP: Cross sections using RDWIA response functions from Udías et al.

- At the moment only a single bin in (ω, q, P<sub>miss</sub>, φ) has been used for each kinematics.
- This will be improved on in the near future, with the goal of binning the data (and the simulation) in physical variables (ω, q, P<sub>miss</sub>, φ).

$$\frac{d^5\sigma}{d\Omega_e d\Omega_p dE_f} = \frac{N_s f_s}{\Delta\Omega_e \Delta\Omega_p \Delta E_f L}$$

where  $f_s$  is the correction to the number of counts, so far being only N5/T5. This will be improved in the future. *L* is our luminosity which takes into account the effective charge, and  $\Delta_e$  is the electro energy range.

## PRELIMINARY Results for <sup>12</sup>C, $P_{miss} = 0$

Q = 0.84(Gev/c), showing MCEEP with ( $E_x(MeV)$ ,final state) = (0,3/2<sup>-</sup>), (2.125,1/2<sup>-</sup>), (5.020, 3/2<sup>-</sup>)



### Achieved Goals: Run 2

#### PRELIMINARY

✓ Improved coincidence time down to 4ns at FWHM. Expecting improvements



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### PRELIMINARY

- ✓ Improved coincidence time down to 4ns at FWHM. Expecting improvements
- $\checkmark\,$  Promising initial raster corrections



### PRELIMINARY ONLINE SPECTRA

• Thin Lead seems a little decent since the start



#### PRELIMINARY

## Run 2 Beginnings

### PRELIMINARY ONLINE SPECTRA

- Thin Lead seems a little decent since the start
- Thick lead will need more work



PRELIMINARY

### PRELIMINARY ONLINE SPECTRA

- Thin Lead seems a little decent since the start
- Thick lead will need more work
- Things should look better in the future

- Complete analysis of Run 2 data
- Theory simulations must be tuned to simulate as much as possible the experimental conditions.
- Radiation in the simulation has to be checked.
- Recheck efficiency corrections of the data to establish final systematic errors
- Compare results from single carbon foil to diamond foil in data
- Explore options to extract carbon spectra from <sup>208</sup>Pb and <sup>209</sup>Bi data