The Hard Photodisintegration of a Proton Pair in ³He Nuclei

I. Pomerantz

R. Gilman, D. Higinbotham, E. Piasetzky, and S. Strauch for the Hall A Collaboration December 15, 2009

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The Process

A process in which a high energy photon is absorbed by a proton pair, resulting in two protons with large transverse momenta.





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Introduction

Experimental setup Analysis E03-101 results Extended run proposal Summary

Motivation

- Significant effort has been devoted to investigate the hard photodisintegration of the Deuteron.
- d(γ,p)n cross section scales with s⁻¹¹ at fixed large cm angle, above 1 GeV in photon energy.
- This scaling is predicted by the Constituent Counting Rule.



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Motivation

Several quark models, have been formulated, each using very different assumptions to explain the high PT in the scaling regime.



I. Pomerantz The Hard Photodisintegration of a Proton Pair in ³He Nuclei

Two possible mechanisms can explain the high PT

1. Breaking a transverse compact object.

- The high PT Results from initial state correlation.
- A very minute part of the pair wave function.



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Two possible mechanisms can explain the high PT

2. Hard re-scattering.

- One proton absorbs the photon, and then interacts with the other member of the pair. The high PT results from this final state interaction.
- Also a rare case (large pp c.m. scattering angle).



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Theoretical predictions





The Hard Photodisintegration of a Proton Pair in ³He Nuclei

Experimental setup

Technique follows that of I Pomerantz et al, E03-101, submitted to PLB.



Kinematic reconstruction

- 6 known variables: $P_1, \phi_1, \theta_1, P_2, \phi_2, \theta_2$.
- 4 unknown variables: $P_n, \phi_n, \theta_n, E_{\gamma}$.
- 4 constraints of energy and momentum conservation:



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$$E_{\gamma} = E_n + E_1 + E_2$$

$$E_{\gamma} = P_1 \cos\theta_1 + P_2 \cos\theta_2 + P_n \cos\theta_n$$

$$P_1 \sin\phi_1 \sin\theta_1 + P_2 \sin\phi_2 \sin\theta_2 + P_n \sin\phi_n \sin\theta_n = 0$$

$$P_1 \cos\phi_1 \sin\theta_1 + P_2 \cos\phi_2 \sin\theta_2 + P_n \cos\phi_n \sin\theta_n = 0$$

Kinematic reconstruction

- Assuming a two-body process by demanding photon energy of a 140 MeV bin off the tip of the bremsstrahlung spectra.
- Since we extract the photon energy for each event, we care very little about beam parameters.



Event selection

• Particle ID.



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- Particle ID.
- Coincidence time, and reaction point.



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Event selection

- Particle ID.
- Coincidence time, and reaction point.
- Phase space cuts on nominal HRSs acceptance.
- Cut on an 140MeV bin in photon energy.

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Electroproduction background subtraction



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Coincidence efficiency

• The edge of the HRS acceptance is set to detect a pair at rest disintegrated in 90° c.m. from a photon with $E_{\gamma} = E_{Beam}$.



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Coincidence efficiency

- The edge of the HRS acceptance is set to detect a pair at rest disintegrated in 90° c.m. from a photon with $E_{\gamma} = E_{Beam}$.
- The Fermi motion of the proton pair and the different photon energies results in a phase-space spread of the outgoing protons larger than the HRS acceptance.
- To correct the cross section calculation for those missed events, we constructed a MCEEP simulation.



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MCEEP results



I. Pomerantz

The Hard Photodisintegration of a Proton Pair in ³He Nuclei

Results

- For $\gamma d \rightarrow pn$, scaled cross section falls continuously above Δ resonance, and scaling starts $\sim 1 \text{ GeV}$ and at $p_T > 1.3$ GeV/c.
- For $\gamma pp \rightarrow pp$, scaled cross section peaks ~ 1.5 GeV, and scaling starts ~ 2 GeV.



Results

- In the scaling region, the cross section ratio $\sigma(\gamma d) / \sigma(\gamma pp) \sim 20$, after correcting for measurements only being up to $p_n = 100 \text{ MeV/c.}$
- HRM explains the low magnitude of the scaled cross section by a cancellation of the opposite sign of the NN helicity amplitudes φ₃ and φ₄.

M. M. Sargsian and C. Granados, Phys. Rev. C80,

014612 (2009)



Energy dependence of the cross section

- Two more data points at $E_{\gamma} = 2.2$ and 4.4 GeV, to look for deviation from constituent counting rules scaling.
 - E03-101
 - Extended measurement (projected)
 - --QGS
 - RNA
 - HRM



α_n distribution

- The light-cone momentum distribution of the spectator neutron, defined as: α_n = <u>E_n-p_n^z</u>/_{m_{3µn}/3}.
- High-energy small-angle final-state rescattering does not change α_n.
- RNA model predicts a much broader distribution of α_n due to selection of large momenta protons from the ³He wave function.



α_n distribution

• We will cover $0.8 < \alpha_n < 1.2$ at 5 kinematics for $E_{\gamma} = 2.2$ GeV.



Kinematics and expected yields

$E_e \approx$	Tgt	α_n	θ_p	Pp	$s^{11} \frac{d\sigma}{dt}$	rate	time	yield
E_{γ}								
[GeV]			[deg]	[GeV/c]	[kb GeV ²⁰]	[cnt/Hr]	Hr	# evts
2.2	³ He	0.8	56.36	1.795	0.001	10	6	54
2.2	³ He	0.9	54.42	1.805	0.007	72	2	144
2.2	³ He	1.0	52.55	1.808	0.02	210	4	840
2.2	³ He	1.1	50.66	1.806	0.004	39	3	117
2.2	³ He	1.2	48.76	1.799	0.001	6	8	46
2.2	d	1.0	52.55	1.808	0.4	8400	1	8400
4.4	d	1.0	42.72	2.990	0.4	21	48	1008
4.4	³ He	1.0	42.72	2.990	0.01	0.45	222	100

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Summary

- We measured hard pp photodisintegration in ³He at $\theta_{c.m.} = 90^{\circ}$, for $E_{\gamma} = 0.8$ to 4.7 GeV.
- For $E_{\gamma} = 1$ 2 GeV, there are structures in pp system not seen in the pn system.
- For $E_{\gamma} > 2$ GeV, $d\sigma/dt \sim s^{-11}$, but γpp cross section is $\sim 20x$ smaller than γpn cross section.
- 15 days of beam time will enable us to determine if deviation from constituent counting rules scaling exists, and to measure α_n to determine if initial or final state effect.
- Looking into adding Hall A neutron detector!