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#### <u>Outline</u>

- □ SBS tracking concept
- □ SBS GEM tracker overview
- □ Back tracker GEM module production Status.
- □ Front tracker GEM module production Status.
- □ GEM readout electronics.
- $\hfill\square$  Conclusion

## SBS Concept

Detectors behind a large dipole magnet:

- Major advantages which pave way to large FOM:
  - □ Large solid angle
  - □ Large momentum bite
  - □ Straight line track analysis.
  - Detectors shielded from charged particle background.
- □ Consequences:
  - □ High rates at detectors.
  - Need good coordinate resolution.

## Why GEMs?

- Super Bigbite Spectrometer concept leads to high rate in trackers: up to 500 kHz per cm<sup>2</sup> in the front detector, and requires good resolution.
- Gas Electron Multiplier (GEM) detectors provide a cost effective solution for high resolution tracking under high rates over large areas.
- Rate capabilities higher than many MHz/cm<sup>2</sup>
- High position resolution ( < 75  $\mu\text{m})$
- Ability to cover very large areas (10s 100s of m<sup>2</sup>) at modest cost.
- Low thickness (~ 0.5% radiation length)
- Already Used for many experiments around the world: COMPASS, Bonus, KLOE, TOTEM, STAR FGT, ALICE TPC, pRAD etc.
- And planed for many future experiments:, CMS upgrade, SoLID, Moller, P2 @ Mainz



GEM foil: 50  $\mu$ m Kapton + few  $\mu$ m copper on both sides with 70  $\mu$ m holes, 140  $\mu$ m pitch



Novel technology: F. Sauli, Nucl. Instrum. Methods A386(1997)531



#### SBS GEM module technical requirements

All GEM modules must be constructed such that:

- all foils have an average dark current of less than 5 nA
  for each 20 x 5 cm<sup>2</sup> sector at 550 V across the foil.
- □ a gain of at least 5000 at the operational voltage in a gas mixture of 70% Argon and 30%  $CO_2$
- a track efficiency of at least 95%, averaged over the module, in cosmic tests
- $\square$  a position resolution of  $\sigma$  < 100  $\mu\text{m}.$
- $\square$  A timing resolution of  $\sigma$  < 25 ns.



### **Current Status of UVa GEM** production

#### **GEM module Production Status**

#### • In the final stretch now

- 37 modules completed.
- 35 modules fully tested:
  - 33 modules 100% operational.
  - 2 modules have one bad sector each. 97% of active area operational.
- Modules #36 and #37 prepared for testing now.
- Module # 38 under construction now
- Moving forward at the rate of 2 chambers per month
- Expected completion of 40 modules by January 2016.
- Five modules: 12-17 moved to Jlab for beam tests
- All material for 8 spare modules have been ordered; expect to complete 8 spare modules by May 2017.



### Path to 40 modules

Module #	Status	
36	Construction complete; prepared for final testing in x-ray box	
37	Construction complete; prepared for final testing in x-ray box	
38	3 GEM foils at UVa, RO at UVA	
39	3 foils and 1 RO to be included in the shipment in the week of Nov. 13	
40	3 foils and 1 RO to be included in the shipment in the week of Nov. 13	

- GEM foils:
  - 147 ordered (need 120 + 24 spares)
  - 127 received and tested: 113 accepted.
    - 20 bad foils with issues (7 repaired: 13 TDB)
    - 1 foils sent back for inspection (re: low gain issue): will be replaced.
- Readout foils:
  - 48 ordered (need 40 + 8 spares)
  - 42 received: 38 accepted so far
  - One with minor issues sent back for repair
  - 3 unacceptable quality (too much Kapton as reported last year); CERN will replace these.

#### **Challenges: Chamber Components**

- Major delay in CERN GEM foil and readout plane production : due to some thickness irregularities found in the raw material foil used for GEM production. (worked closely with Dick Majka and colleagues to identify problem)
- We had some foils in reserve; however in many occasions have had to wait for the GEM foils and readout.
- Sent back a batch of questionable foils for repair; some were fully repaired.
- Discovered a previously unobserved issue with 6 foils: positive current spikes with arc discharges around GEM holes: have not been fixed after repair;
- Major and unfortunate issue with the bath received on 11/01
- All problem foils and 1 RO shipped back to CERN on 11/03





## Challenges

- ~ 1 month unexpected delay due to frame coating varnish going bad.
  - Specific Varnish from a Swiss company, certified by CERN re: chamber aging.
  - Hazardous liquid; so shipping, customs clearance takes 1 month (and ~ \$2k).
  - Arrived in early October
- Chinese visiting student left in June: production slowed down while new student was coming up to speed.
- Will complete all 40 modules and couple of spares by project end date; but need to figure out manpower for ~ 6 spare modules .

- Currently ongoing
- Five modules separated by 10 cm each: very similar to a SBS tracker.
- Triggered by a lead-glass matrix at center
- Goals: Identify good tracks in a high rate background, study effectiveness of timing and charge correlation cuts to suppress background
- Currently at 70°, occupancy is ~ 1.2%; going to a smaller angle this week to get higher













### Work in Progress: GEM holding frame



#### First Holding frame completed at Uva:

- Everything fits really well !
- Minor changes in the design to strengthen corners to improve rigidity.
- Ten frames of modified design in production now

#### Front Tracker Status

#### FT - Production Summary and Plan

- 14 GEM modules produced, 2 under completition (out of 18 by end of 2016, 4 spare modules expected by spring 2017)
- 13 tested or under testing, 1 damaged during assembling, 3 with one GEM sector disconnected (work on fixing it)
- 2 full chambers assembled at JLab, need to be re-tested
- Improved cosmic test stand in Italy (top)
- Cosmic test stand at JLab under finalization (bottom)
- GEM readout integration in CODA/DAQ almost completed

One year plan:

- Complete GEM modules assembling (including spares)
- Extended chamber cosmic test
- Complete data suppression in hardware and finalize integration in CODA/DAQ
- Implement robust tracking for FT





## Second GEM Chamber at JLab (Nov 2015)



5 carbon frame structure under construction; next chamber expected end of 2016

## Study of robust FT Track Reconstruction

- Multistep approach:
  - Hit association: Neural Network
    - (need smart energy function)
  - Precise tracking: Kalman filter
    (rather consolidated approach)
- (Slow) work in progress:
  - Consolidate NN
  - Implementation of artificial
    - **RETINA** approach
    - (for hit association)

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# SBS Front Tracker - Workforce and funding

INFN Group	Researcher (Unit / FTE)	Tech. (Unit / FTE)	Role	
Bari	1/0.3		Gas system and beam test	
Catania	3* / 1.6	3 / 1.5	GEM module assembling, mechanics, beam test, analysis	
Genova	1**/0.3		Electronics design and test	
Rome/Sanità	2 / 0.6	3 / 1.5	Coordination, design, test, chamber integration, analysis, DAQ and track reconstruction	
Total	7/2.8	6/3		
Other support: CERN / IIVa / II ab				

- \*) one PhD student
- \*\*) electronic engineer
- INFN/Funding:
  - Prototyping and Production (2008-2016): 900 kUSD
  - Commissioning and Maintainance (2017-2019): ~40 kUSD/year

**GEM Readout Electronics** 

## GEM – APV-MPD based Readout Electronics



- 128 analog ch / APV25 ASIC
- 3.4 µs trigger latency (analog pipeline)
- Capable of sampling signal at 40 MHz
- Multiplexed analog output (100 kHz readout rate)
- Up to 16 APV25 cards (2048 chs) on a single MPD (parallel readout)
- Altera Arriga GX FPGA / RAM: DDR2 (128 MB)
- Optical Fiber Link interface (Aurora ~2 Gb/s peak)
- 100 MHz system clock and Front panel coax clock
- Used HDMI-A for analog and digital signals
- VME/32, VME64, VME64-VXS compliant (up to 200 MB/s peak)
- 4 high speed line on the VXS available for data transfer
- Firmware v. 4.0 (74% resources):
  - Finite-Impulse-Response Filter (16 parameters)
  - Zero Suppression (sparse readout)
  - Common mode and pedestal subtraction
  - Remore reconfiguration
  - ~2 ns trigger time resolution
  - VME / Optical Fiber simultaneous implementation



#### MPD Firmware latest status

- Version 4.0 74 % FPGA resources used
- New (optimized and better organized) memory map tested succesfully
- All VME cycles tested; minor issue to be fixed
- Optical Link to SSP under deep testing (thanks to the CODA setup with SSP and Intel CPU recently installed in Italy)



#### Back Tracker Electronics / as of Nov 2016

- 57 MPD ordered\*: 57 delivered and tested
  - 3 MPD did not pass the test and sent back for repair: expected to befixable
- 925 APV cards ordered/ 927 delivered/880 needed
  - Testbench used in Italy with automated procedure mainly to identify bonding quality by SNR measurement; was able to test 50 to 80 cards/day. Each card has a pdf test report. A summary report is also generated
  - 147 tested so far; 7 with issues
- Received all backplanes: 42 long/82 short (40 short/80 long needed)
- All LV regulator boards and patch panels fabricated







\*Only 57 MPD are required to cover all channels; however optimum cabling arrangement requires 70 MPD. Plan to purchase 13 more MPD (+2 spares) using UVa funds 27

- □ The SBS back tracker production is going well
  - $\Box$  Now in the final stretch: 37 modules and counting.
  - □ Expect to complete 42 by end of January.
- □ Beam test with 5-module telescope currently going on
- Front tracker: 9/18 modules produced and one full chamber assembled.
- □ Most of the MPD electronics produced in Italy and received at UVa.
- □ Integration of GEM readout into CODA currently underway

Backup Slides



## SBS FT – Tracking Cosmic (preliminary)

Tree GEM modules

- Choose one GEM as «test» module and assume the other two as «calibrated»
- 2. Select «true» hits on the two «calibrated» modules
- 3. Compute the straight line passing for the two hits
- 4. Project the line on the third «test» module
- 5. Compute the 2D distance of the projected point from the measured hit (right)
- 6. Do 1. for each module

Raw manual alignment with accuracy at the level mm on each axis.

Measured 2D residues at the level of 0.3 mm:

- deconvoluting the error on each chamber: ≈1/sqrt(3)
- From 2D to single axis: ≈1/sqrt(2)
- Error on single axis: ≈ 0.3/sqrt(3\*2) ≈ **0.12 mm**

Cosmic data taking with up to 7 modules in progress



2D distance between projected and measured points

### MPD for SBS-BT-GEMs: Large production of MPD electronics

# 40 modules for the 2 Back Tracker stations (112,640) electronic channels to readout

- Original option: combine two detector readout strips into one electronic channels
- A lot of concerns for this option:
  - Additional items: Need adapter 2-1 strips in addition to the back planes
  - electronic noise level, APV25 saturation, detector performance ...
- Second option: Reading out all single detector strips
  - •Safer option for performance and less development needed
- Economies of scale and strong \$ allow us to instrument all channels instead: big improvement.
- However: forcing us to really stretch the available resources.
- Including spares procuring 118 k MPD channels for the back tracker.
- Order for (almost) all components placed.
- 57 MPD units received and tested: two are bad, sent back for repair.
- All APV cards received; testing ongoing
- All back planes received.

#### APV25 - Long (23 m) cable effects on analog signals

Digital

• Problem:

NO-FIR

FIR-8

**FIR-12** 

- The large «binary» information (digital header) at the beginning of the analog signals of the APV introduce a large noise on the first (~20) channels of the frame
- Longer the cable larger the noise, higher the numbe of channel involved
- Belle (2012 JINST 7 C01082) proposed a 8-parameter FIR filter (12 m long cables) in firmware
- We adopted: two different FIR implementations in firmware (one tested) and added an off-line pedestal subtraction dependant on the digital header value (LUT suppression): ⇒ very noisy channels largely recovered

Avg: 12.5

Max: ~100

Avg: 13.6

Max: ~60

Avg: 11.7

Max: ~45

Avg: 9.1

Max:~40

Avg: 11.1

Max: ~60

Avg: 11.3

Max: ~50



128 Analog

Ticks