

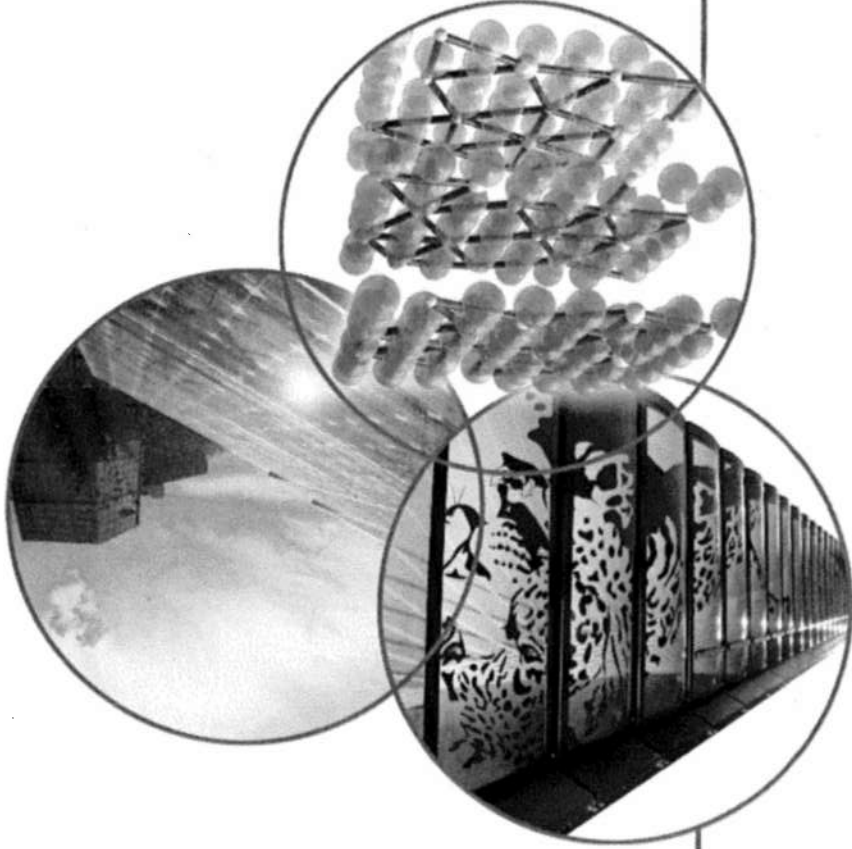
Neutron Matter as a Composite Bose- Fermi Superfluid

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Overview

• The proposed Hamiltonian model combines:

1. An NN interaction based on three- and six- quark clusters

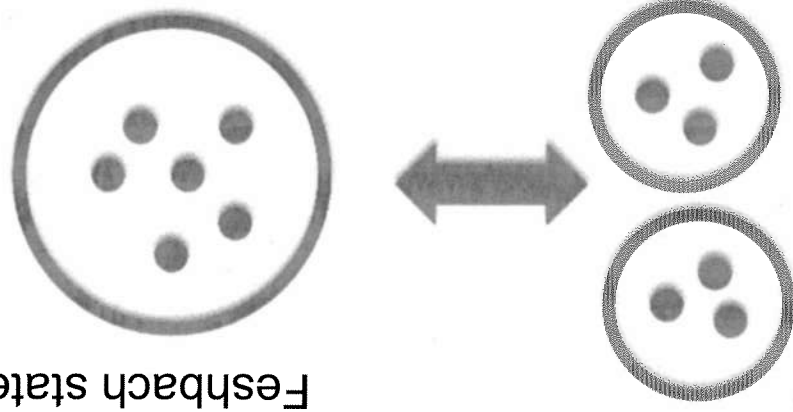
• Nucleon = cluster of three quarks,

• Two nucleons couple to a cluster of six-quarks

• The six-quark cluster short-lived non-resonant Feshbach state(s)

• Feshbach state energy and form factor are found from the NN phase shifts

• A single Feshbach state may be sufficient to account for $E_{lab} > 350 \text{ MeV}$



Feshbach state (six quark cluster)

2. Many-body model of composite Bose-Fermi superfluid leverages the atomic physics model for the same.

Two-body interaction by coupling nucleons to Feshbach state(s)

1) NN interaction via Feshbach State(s)

- General Feshbach projection operator method yields:

$$\tan \delta(k) = -k K_{PP}(k) / (4\pi)$$

$$K_{PP}(k) = \frac{H_{PQ} E - H_{QQ} - \text{P.V.} \frac{H_{QP} E - H_{PP}}{1}}{H_{QP}}$$

- Allows for any number of Feshbach States in Q-space

- For a single Feshbach state:

$$H_{QQ} = k_1^2 \quad \text{in units } \hbar^2/m$$

– Phase shift changes sign at k_0

$$H_{PQ} = \sqrt{4\pi} f(k)$$

$$f(q) = (q^2 - k_0^2) h(q)$$

$$h(q) = g e^{-q^2/s^2}$$

← F. Tabakin, Phys. Rev. C 174, 1208 (1968);
Kukulin et al., Phys. of Atom. Nuclei, 64, 1748, (2001)

Phase-shift for a single Feshbach state

- Phase-shift changes sign k_0 :

$$\tan \delta(k) = -k \frac{k_2 - k_2^1 + \int_0^{\frac{\pi}{2}} f_2(q) \frac{q^{k_2 - k_2^1}}{b_2(q)} dq}{f_2(k)}$$

$$\tan \delta(k) = -k \frac{1 + \int_0^{\frac{\pi}{2}} q^{k_2 - k_2^1} h_2(q) \frac{q^{k_2 - k_2^1}}{b_2(q)} dq}{(k_2 - k_2^1) h_2(k)}$$

$$k_2^1 = k_2^0 + \frac{\pi}{2} \int_0^{\infty} (q^2 - k_2^0) h_2(q) q^2 dq$$

- The energy of the Feshbach state $(k_1)^2$ is uniquely determined by k_0 and $h(q)$ in this model.

Fitting Low-Energy 1S_0 nn Data

- Effective range expansion of $\tan(\delta)$ yields expressions for the scattering length and the effective range that are equated to experimental values

$$a = -18.9 \text{ fm}$$

$$r_0 = 2.75 \text{ fm}$$

Form-factor parameters are found using Mathematica:

$$k_1 = 1.62099 \text{ fm}^{-1}$$

$$g = 0.812091 \text{ fm}^{3/2}$$

$$s = 1.44185 \text{ fm}^{-1}$$

for a form factor:

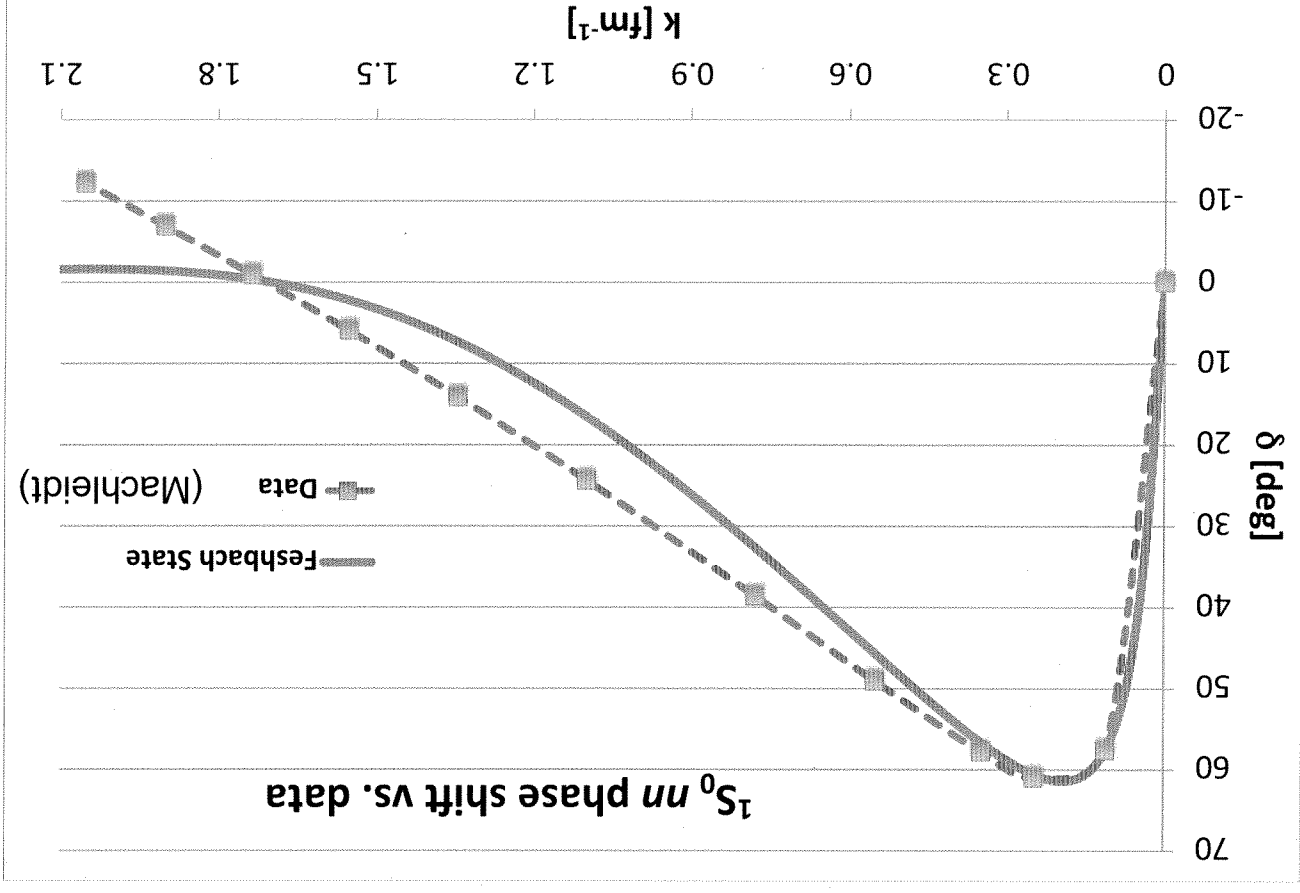
$$f(q) = (q^2 - k_0^2)h(q)$$

$$h(q) = g e^{-q^2/s^2}$$

(The same method was used to find parameters that yield infinite scattering length...)

Comparison to 1S_0 nn phase shift data

- Good fit at low energies *and* it crosses zero at $E_{lab} = 240$ MeV
- An improvement over fitting just (a, r_0)
Schwenk and Pethick,
PRL 95, 160401 (2005)
- suitable for higher densities in a many-body problem
- Additional parameters could improve the fit at other energies



2) Neutron matter Grand Canon. Hamilton.

- Two neutrons couple to Feshbach state(s)

– Simplifies the computation of the interaction (quartic $\rightarrow$ cubic)

$$\hat{K} = \hat{H} - \mu N$$

$$= \sum_{\sigma} \int_{\mathbf{q}} (k^2 - 2\mu) \hat{\psi}_{\dagger}^{\sigma}(\mathbf{q}) \hat{\psi}_{\sigma}(\mathbf{q}) + 2(\epsilon - 2\mu) \int_{\mathbf{q}} \hat{\chi}_{\dagger}(\mathbf{q}) \hat{\chi}(\mathbf{q})$$

$$+ \frac{1}{2} \int_{\mathbf{q}''} \int_{\mathbf{q}'} \hat{\chi}_{\dagger}(\mathbf{q}'') F(\mathbf{q}'', \mathbf{q}, \mathbf{q}') [\hat{\psi}_{\dagger}(\mathbf{q}) \hat{\psi}_{\uparrow}(-\mathbf{q}') - \hat{\psi}_{\dagger}(-\mathbf{q}) \hat{\psi}_{\uparrow}(\mathbf{q}')] + \text{H.c.},$$

A two-body form-factor

where

$$F(\mathbf{q}'', \mathbf{q}, \mathbf{q}') = 2\sqrt{4\pi} f\left(\frac{\mathbf{q}''}{\mathbf{q} + \mathbf{q}'}\right) \delta(\mathbf{q}'' - (\mathbf{q} - \mathbf{q}'))$$

- A mean molecular field approx. $\langle 0 | \hat{\chi}(\mathbf{q}) | 0 \rangle = \delta(\mathbf{q}) \frac{\chi}{\sqrt{2}}$ (cubic $\rightarrow$ quadratic):

$$\hat{K} = \sum_{\sigma} \int_{\mathbf{q}} (k^2 - \mu) \hat{\psi}_{\dagger}^{\sigma}(\mathbf{q}) \hat{\psi}_{\sigma}(\mathbf{q}) - (2k_2^2 - 2\mu) \chi^2$$

$$+ 2\sqrt{4\pi} \chi \int_{\mathbf{q}} f(\mathbf{q}) \hat{\psi}_{\dagger}(\mathbf{q}) \hat{\psi}_{\uparrow}(-\mathbf{q}) + \text{H.c.}$$

The First Law of Thermodynamics

- Use it to relate pressure, energy density, and particle density

$$dE = -PdV + \mu dN$$

$$P = -\mathcal{K}/V = -n + \mu \rho$$

$$e = \frac{d}{n} = \mu - \frac{d}{\rho}$$

Neutron matter gap function and EOS

- Application of the Bogoliubov transform technique and the First Law of thermodynamics yields:

$$P = 2 \int^q E(q) v_2^2(q) - 2(k_1^2 - \mu) \chi_2^2$$

$$\rho = 2 \int^q v_2^2(q) + 2\chi_2^2$$

- Two fermions per boson
 - Energy units in the lab frame $\hbar^2/2m$

where

$$E(q) = \sqrt{(q^2 - \mu)^2 + 16\pi\chi_2^2 f_2^2(q)}$$

$$v_2(q) = \frac{1}{2} \left[1 - \frac{q^2}{q^2 - \mu} \frac{E(q)}{f_2(q)} \right],$$

- Extremizing pressure w.r.t. mean field yields a gap equation

Equation of State

- Extremizing pressure w.r.t. the mean field yields a gap eq.:

$$\frac{\partial \chi}{\partial P} = 0 \Leftrightarrow \mu - k_1^2 + \frac{\pi}{2} \int_0^\infty \frac{f_2(q) E(q)}{q^2} dq = 0$$

- For a given μ this integral equation is solved for χ
- the energy per particle is $e = u/p = \mu - P/p$

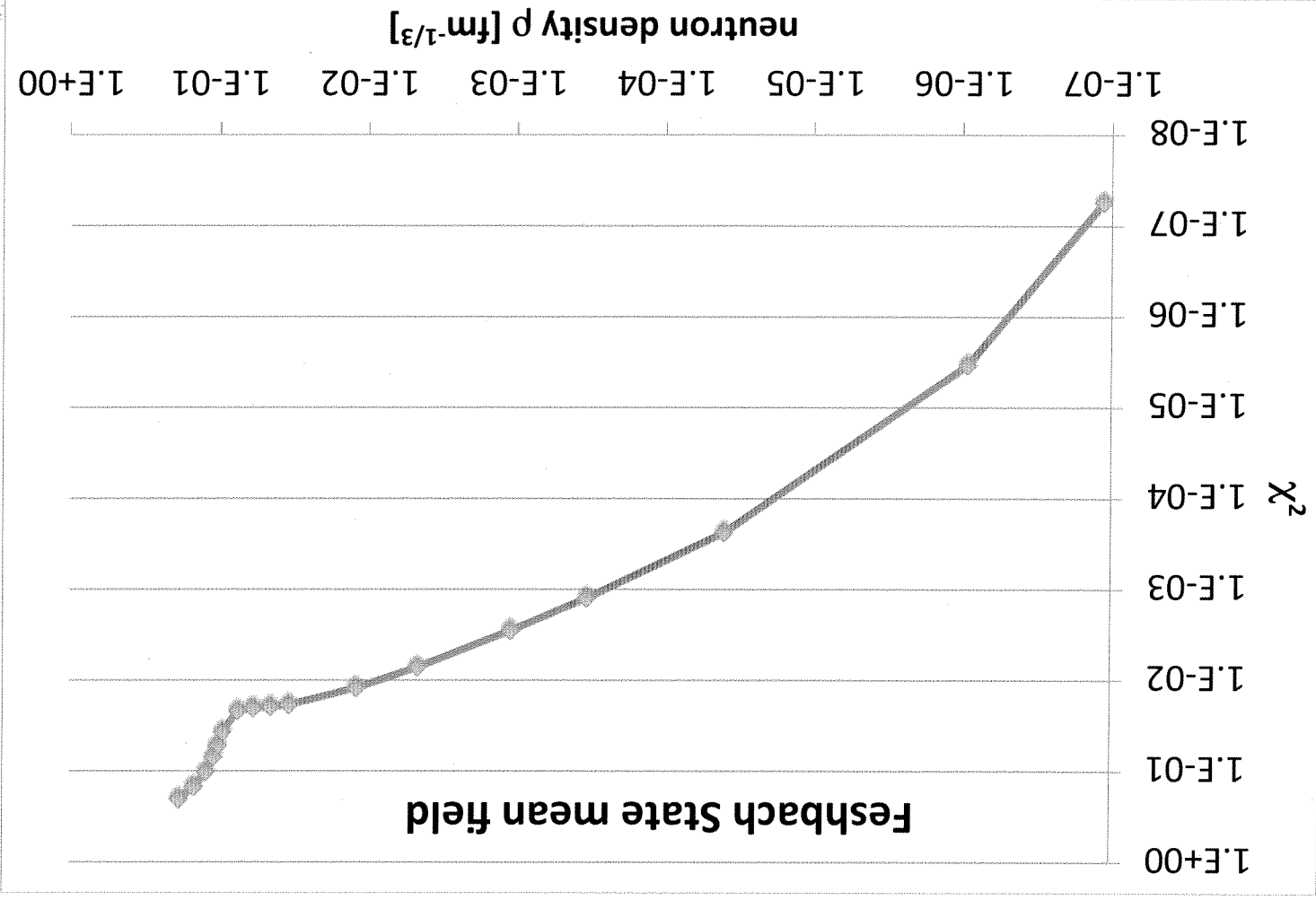
that is plotted in units of energy per particle of a Fermi gas:

$$k_2^F = (3\pi^2 \rho)^{2/3} \neq \mu$$

$$e_{FG} = \frac{3}{5} k_F^2$$

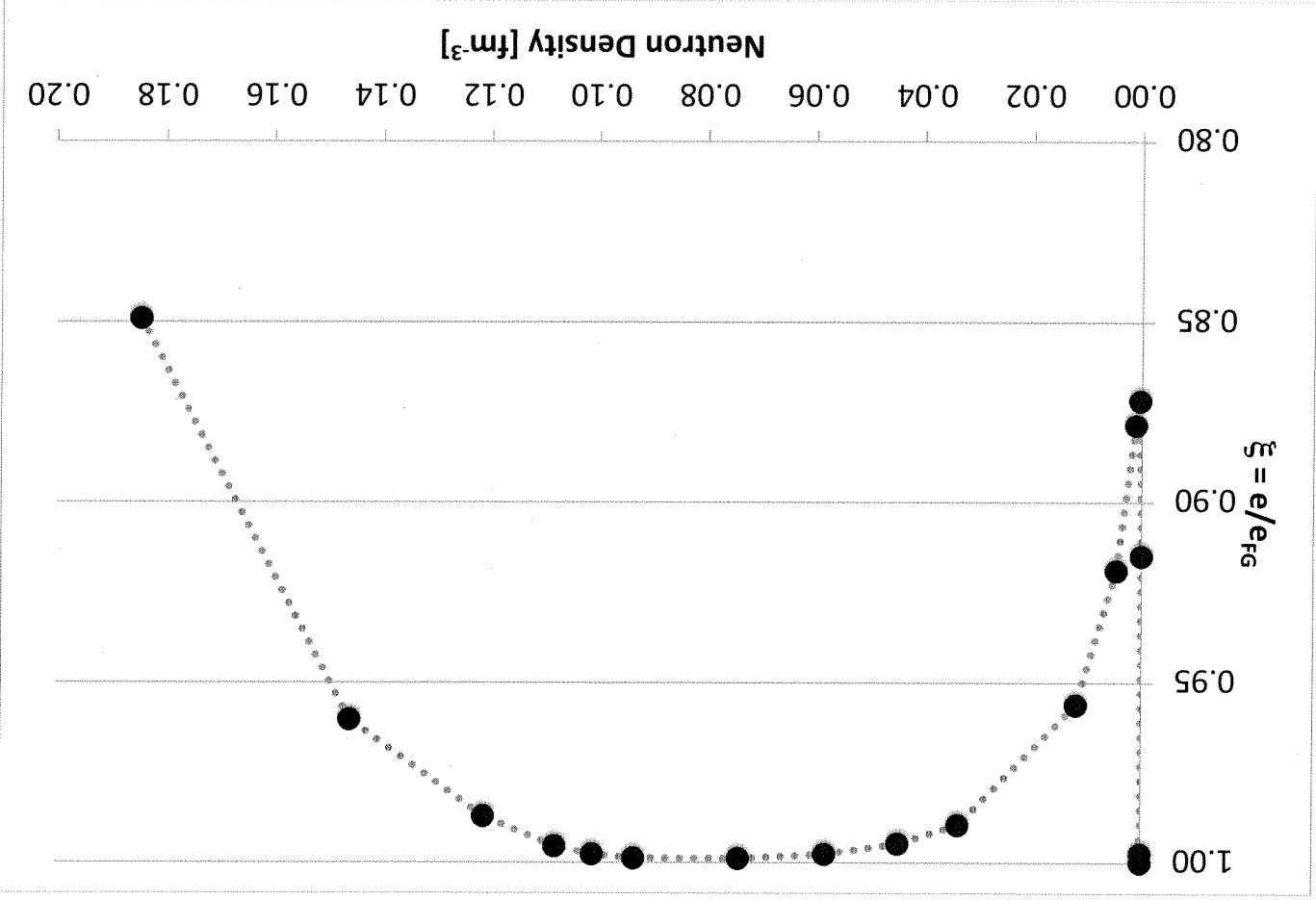
Mean field as a function of neutron density

- A solution to the gap equation



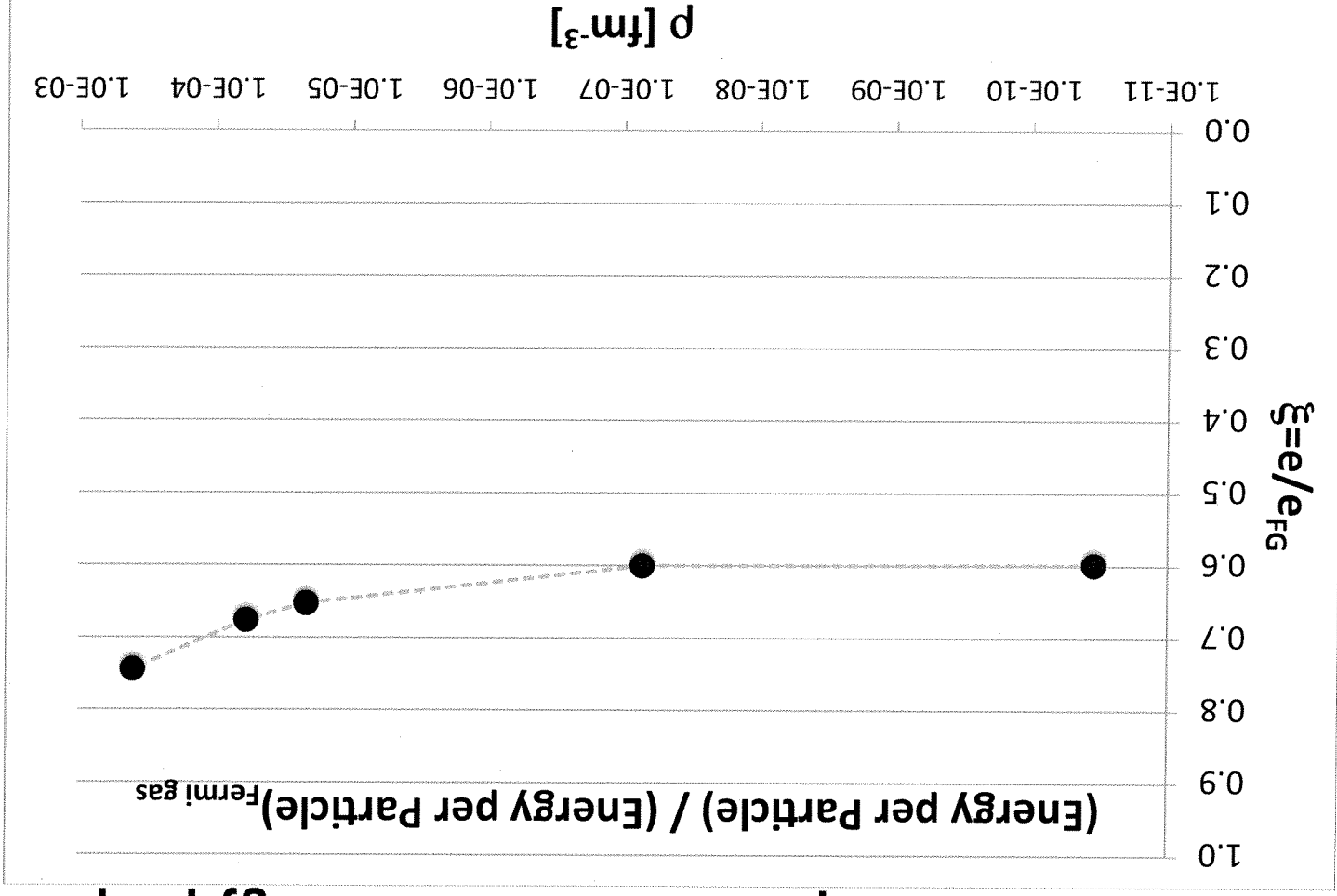
Neutron Matter EOS

- Energy per particle is higher than in extant models
 - However, it is always smaller than that of a free Fermi gas.
- This computation does not include RPA correction
 - RPA correction are expected to lower the energy per particle



Unitary Fermi Gas EOS

- Energy per particle is higher than in extant models — However, it is always smaller than that of a free Fermi gas.
- This computation does not include RPA correction



RPA correction are expected to lower the energy per particle

Summary and Outlook

- A novel use of a six-quark cluster NN interaction in a composite bose-fermion field theory at $T=0$ is suggested
 - Feshbach state parameters were fitted to 1S_0 phase shift
 - EOS was solved for neutron matter and a unitary Fermi gas
 - Energy per particle was found to be higher than in extant models
 - RPA correction is needed to decrease energy/particle (in progress)
- Outlook:
 - Extend to symmetric nuclear matter
 - Apply to ground states of nuclei
 - Apply to excited states ($T>0$)
 - Connect to nuclear reactions?