

Unpolarized Radiative Corrections

Interpolation Systematics and Unfolding Methods

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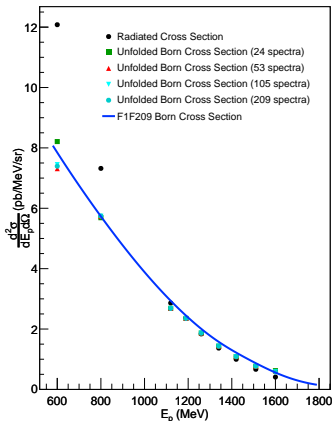
Outline

- 1 Interpolation Systematics
- 2 Unfolding Methods
- 3 Summary

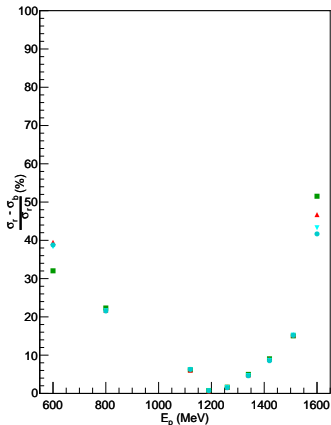
Interpolation Systematics (1)

4-pass Data

Radiative Corrections at 4-pass ($E_e = 4730$ MeV, $\theta = 45^\circ$)



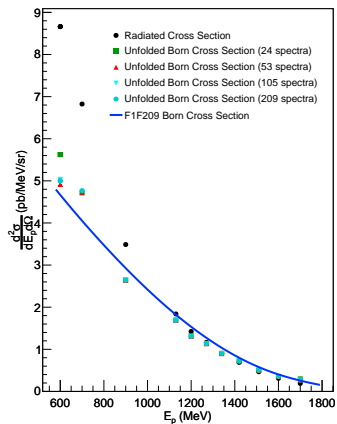
Size of Radiative Corrections



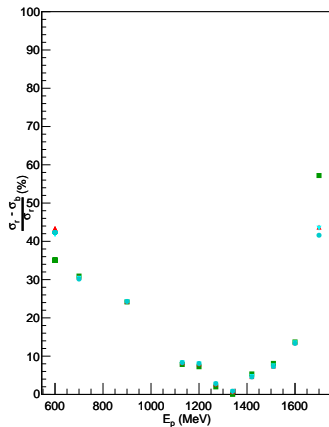
Interpolation Systematics (2)

5-pass Data

Radiative Corrections at 5-pass ($E_e = 5890$ MeV, $\theta = 45^\circ$)



Size of Radiative Corrections



Comparison of the Two Methods (1)

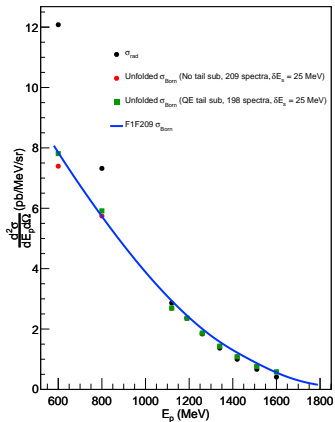
Description

- Review of methods:
 - 1 Subtract the QE tail (as calculated in ROSETAIL) and then unfold from the **pion** threshold
 - 2 Unfold from the **elastic** threshold
- Given how the unfolded results change with the number of spectra used in the interpolation, what happens if we change the number of spectra when integrating from the π threshold?
 - ▶ Use the same spacing in the E_s integral as seen in the elastic threshold results for 209 spectra
- In theory, we **should** get the same result. . .

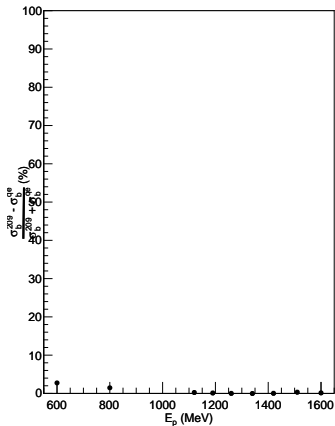
Comparison of the Two Methods (2)

4-pass Data

Radiative Corrections at 4-pass ($E_s = 4730$ MeV, $\theta = 45^\circ$)



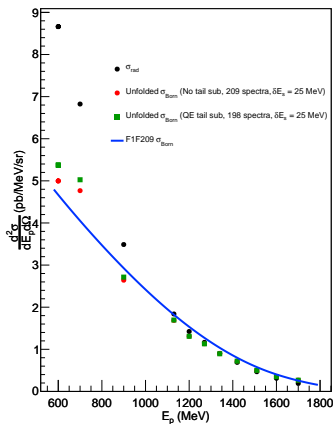
Percent Difference



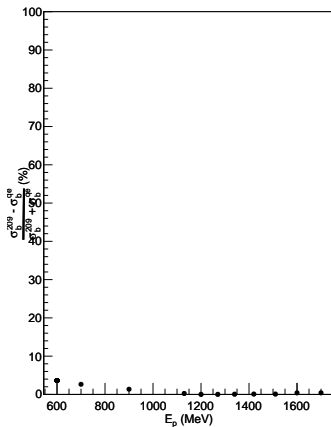
Comparison of the Two Methods (3)

5-pass Data

Radiative Corrections at 5-pass ($E_e = 5890$ MeV, $\theta = 45^\circ$)



Percent Difference



Summary and What's Next

- Unpolarized Radiative Corrections
 - ▶ Changing the number of spectra has an effect in the unfolding, tends to converge after ~ 50 spectra. . .
 - ▶ The two methods of unfolding agree for most bins in E_p , with disagreement at the $\sim 3\text{--}4\%$ level at the lowest bins in E_p
 - ★ Model dependence?
- Polarized Radiative Corrections
 - ▶ Getting closer on the C++ POLRAD version
 - ▶ Fortran version still needs work. . .
- What's next?
 - ▶ Investigate systematics of RADCOR interpolation method
 - ★ Model dependence
 - ▶ Continue work on POLRAD (C++/Fortran)