

# E06-014 Analysis Update Radiative Corrections and $A_1^n$

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# Outline

- 1 Radiative Corrections Update
- 2  $A_1^n$  Extraction

# Radiative Corrections Update (1)

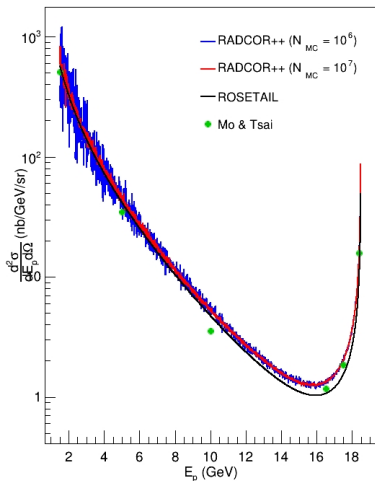
Details: Integration Method

- Moved to Monte Carlo integration to speed things up
  - ▶ Method: Sample Mean (so far)
  - ▶ External radiation: 4D integral would not finish over many days of running; will be 5D for inelastic tail
  - ▶ With MC, can finish a full spectrum in a matter of a few hours (at  $10^7$  MC events)
- In the middle of tuning the importance sampling, accept/reject methods (should make things even faster)

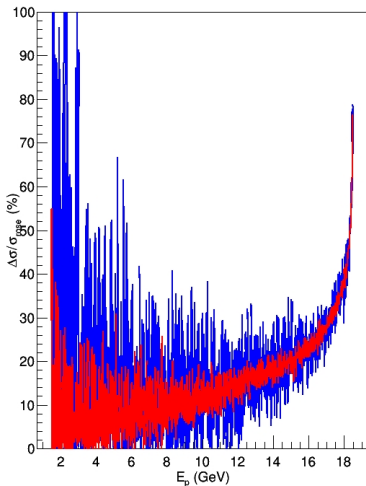
# Radiative Corrections Update (2)

Example Run/Test: Internal Elastic Tail on Proton

Proton Internal Elastic Tail ( $E = 20$  GeV,  $\theta = 5$  deg)



Relative Error of RADCOR w.r.t. ROSETAIL



# $A_1^n$ : First Calculation (1)

## Method and Details

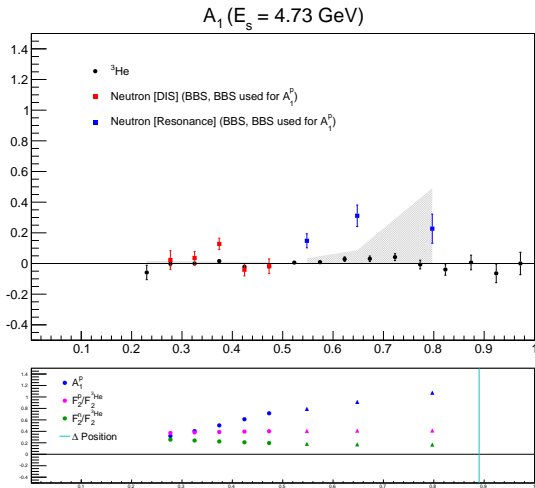
- Utilize the effective polarization method:

$$A_1^n = \frac{F_2^{3\text{He}}}{\tilde{P}_n F_2^n} \left[ A_1^{3\text{He}} - \tilde{P}_p \left( \frac{F_2^p}{F_2^{3\text{He}}} \right) A_1^p \right]$$

- Assume duality and re-bin resonance data into 3 (4-pass) and 2 bins (5-pass)
- $\tilde{P}_i$  terms include off-shell effects:
  - $\tilde{P}_n = 0.879 + 0.056$
  - $\tilde{P}_p = 2(-0.021) - 0.014$
- Used a number of model combinations (unpolarized, polarized) for systematic studies
  - Unpolarized: BBS, CTEQ6
  - Polarized: BBS, BB, DNS, GS, DSSV, LSS
  - Data shown in slides uses BBS for both unpolarized and polarized models

# $A_1^n$ : First Calculation (2)

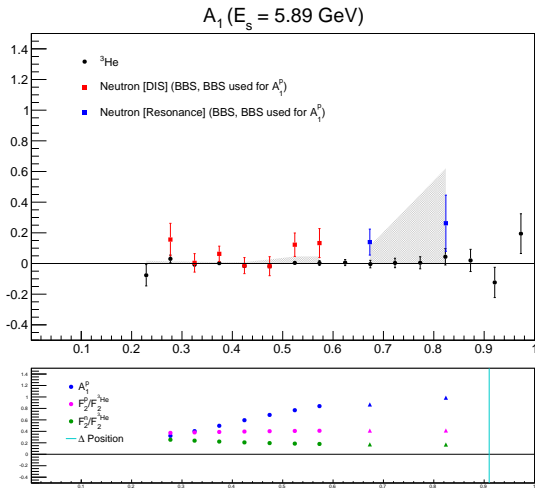
Extracting  $A_1^n$ :  $E = 4.73$  GeV Data



● Grey band shows systematic error due to model dependence

# $A_1^n$ : First Calculation (3)

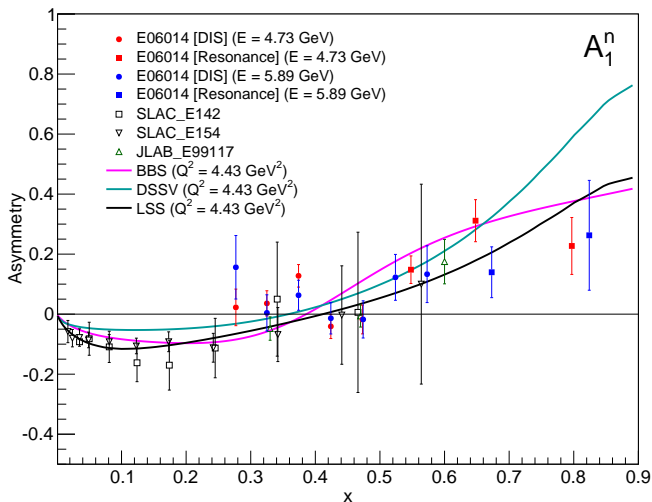
Extracting  $A_1^n$ :  $E = 5.89$  GeV Data



● Grey band shows systematic error due to model dependence

# $A_1^n$ : First Calculation (4)

Compared to World Data





- Radiative Corrections

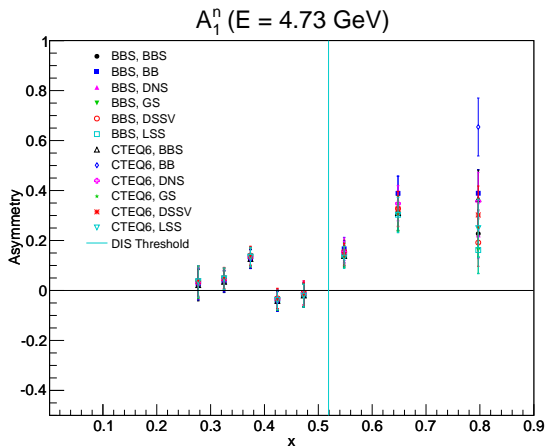
- ▶ MC integration techniques seems to be the best way to go with multi-dimensional integration
- ▶ Fine-tune importance sampling, accept/reject methods (in  $d$ -dimensions)
- ▶ Finish debugging POLRAD++

- $A_1^n$  Extraction

- ▶ Computed  $A_1^n$  using the effective polarization method
- ▶ Utilizing different combinations of input models, model dependence is found to be small in the DIS region
- ▶ Consistent with world data, but precision is not as good as some other experiments (e.g., E99-117)

# Backup (1)

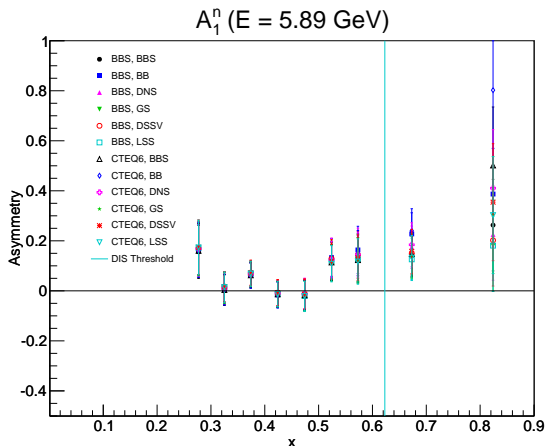
$A_1^n$ : Model Dependence for  $E = 4.73$  GeV Data



- Error taken as the **difference** between the largest and smallest values per bin

# Backup (2)

$A_1^n$ : Model Dependence for  $E = 5.89$  GeV Data



- Error taken as the **difference** between the largest and smallest values per bin